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## Effects of Energy Resolution and Bin Size on the Measurements of Nuclear Neutron Distributions Using Coherent Elastic Neutrino-nucleus Scattering

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# Effects of Energy Resolution and Bin Size on the Measurements of Nuclear Neutron Distributions Using Coherent Elastic Neutrino-nucleus Scattering

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May, 2020

# Abstract

Coherent elastic neutrino-nucleus scattering (CE $\nu$ NS) is a process that involves the neutral-current scattering of a neutrino with an entire nucleus, which probes more bulk properties of the nucleus. Using the Taylor expansion of the form factors and calculation of effective moments that sums over isotopes, we successfully simulate the events for coherent elastic neutrino-nucleus scattering for detectors using Ar and Ge and demonstrate that CE $\nu$ NS can determine both the second moment and the fourth moment. Running the Monte Carlo simulation using a detector filled with 10 tonnes of natural germanium, a neutrino flux of  $3 \times 10^7$  neutrinos per second per  $\text{cm}^2$ , and a bin size of 5 keV, we demonstrate that the neutron radii can be determined with less than 1% uncertainty if the detector efficiency remains constant and that a lower energy threshold can lead to a more accurate result.

## Acknowledgements

I would like to thank Professor Kelly Patton for her ongoing guidance and assistance on this research project throughout the entire academic year, especially during this difficult time period of the COVID-19 pandemic. Furthermore, I would like to thank the Department of Physics and Astronomy at Colby College for this precious opportunity to conduct research and draft the honors thesis for my senior year. Last, I would like to thank my friends and my parents for their encouragement and support that help me stay strong and positive under the COVID-19 situation and make this thesis a reality.

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# 1 Introduction

The neutron radius and the proton radius are two of the most fundamental and important properties of a nucleus. Since protons carry positive charges, the proton charge radii can be obtained with accuracy through electron scattering. However, the uncertainties in calculations and the discrepancies between different theoretical fit models make it hard for us to determine the neutron radii for nuclei with medium to heavy weight.[1] Since the neutron and proton radii are related to the neutron and proton density distributions in nuclei, we can obtain the neutron radii with confidence by measuring the neutron density distributions accurately. Determining neutron radii with precision can generate significant impacts on many related fields, such as astrophysics and experimental particle physics, and provide inspirations for the development of new physics.

## 1.1 Energy Density Functional Theory (DFT)

The most common method for determining neutron densities is the energy density functional theory (DFT).[2] In the Skyrme Hartree-Fock model, the total binding energy is given by

$$E = E_{kin} + \int d^3r \epsilon_{Sk} + E_{Coul} + E_{pair} - E_{corr}, \quad (1)$$

which is the sum of the kinetic energy, the Skyrme energy functional that includes the interactions between nucleons, the Coulomb energy, the pairing energy, and the energy corrections. [3] The Skyrme energy functional is given by

$$\epsilon_{Sk} = \sum_{T=0,1} (\epsilon_T^{even} + \epsilon_T^{odd}), \quad (2)$$

where the odd part contains the time-odd densities and the even part contains the time-even densities that sum over the isospin  $T$ . Ref. [3] contains a detailed description of the mean-field models for nuclear structures. By fitting the nuclear observables obtained from experiments to the various parameters in the model, we can predict the neutron density distribution inside a nucleus.

One interesting observable is the neutron halos for light nuclei with abundant neutrons, which is caused by the weak binding of the outermost neutrons.[3] Although

there are vast disagreements in the parametrization of the halos, we can perform a robust measurement by implementing the Skyrme interactions and measuring parameters such as the root-mean-squared radius, the diffraction radius, and the surface thickness.[3] Another interesting observable is the neutron skin, which can help us determine the difference between neutron and proton radii by determining the relation between the neutron skin and the isovector forces.[3] The asymmetry energy from the interactions predicted by the Skyrme Hartee-Fock model forms a special relation with the neutron skin with higher asymmetry energy producing a larger skin. Therefore, the Skyrme Hartee-Fock model can help predict neutron radii and accurate measurements of neutron radii can help revise the model and improve its predictions.

## 1.2 Neutrino-Nucleus Coherent Scattering

Coherent elastic neutrino-nucleus scattering(CE $\nu$ NS) is a process that involves the neutral-current scattering of a neutrino with an entire nucleus, which probes more bulk properties of the nucleus. The idea of probing the neutron form factor by measuring the neutrino-nucleus coherent scattering events to determine neutron density distributions is first proposed in Ref. [1] using a detector filled with one tonne of  $^{40}\text{Ar}$ . Traditionally, the neutrino-nucleus coherent elastic scattering events have been difficult to detect since the recoil energies are very small. The recoil energies are mostly in the sub-MeV range, which is below the threshold for most neutrino detectors. However, the recent developments of the ultra low threshold detectors enable the detection of weakly interacting massive particles and low energy solar neutrinos with thresholds of 10 keV or even lower. [4] We will introduce some of the different types of these detectors afterwards. These detectors allow us to measure the counts for the neutrino-nucleus coherent scattering and determine the neutron form factors. Apart from the neutron form factors which are our main focus, this section will continue to introduce briefly the other aspects of neutrinos and nuclear physics behind the coherent scattering that can be explored using CE $\nu$ NS.

The rate of the neutrino-nucleus coherent elastic scattering is predicted by the standard model, which is proportional to the weak charge  $Q_W$ . The weak charge is defined as

$$Q_W = N - (1 - 4 \sin^2 \theta_W)Z \quad (3)$$

where  $N$  is the number of neutrons,  $Z$  is the number of protons, and  $\theta_W$  is the weak

mixing angle. If we can measure the CE $\nu$ NS events under a 10% uncertainty, then we can extract the mixing angle within a 5% uncertainty. This provides a new method to determine the weak mixing angle if we can measure the scattering events accurately.

The non-standard interactions of neutrinos also impact the cross section. We can assume an effective Lagrangian for interactions between neutrinos and hadrons as[4]

$$L_{\nu H}^{NSI} = -\frac{G_F}{\sqrt{2}} \sum_{\substack{q=u,d,\alpha \\ \beta=e,\mu,\tau}} [\bar{\nu}_\alpha \gamma^\mu (1 - \gamma^5) \nu_\beta] \times (\epsilon_{\alpha\beta}^{qL} [\bar{q} \gamma_\mu (1 - \gamma^5) q] + \epsilon_{\alpha\beta}^{qR} [\bar{q} \gamma_\mu (1 + \gamma^5) q]). \quad (4)$$

The cross section for coherent elastic scattering of neutrinos of flavor  $\alpha$  is [4]

$$\frac{d\sigma}{dE}_{\nu\alpha A} = \frac{G_F^2}{\pi} F^2(2ME) \left[ 1 - \frac{ME}{2k^2} \right] \times \{ [Z(g_V^p + 2\epsilon_{\alpha\alpha}^{\mu V} + \epsilon_{\alpha\alpha}^{dV}) + N(g_V^n + \epsilon_{\alpha\alpha}^{\mu V} + 2\epsilon_{\alpha\alpha}^{dV})]^2 + \sum_{\alpha \neq \beta} [Z(2\epsilon_{\alpha\beta}^{\mu V} + \epsilon_{\alpha\beta}^{dV}) + N(\epsilon_{\alpha\beta}^{\mu V} + 2\epsilon_{\alpha\beta}^{dV})]^2 \}, \quad (5)$$

where  $g_V^p = (\frac{1}{2} - 2\sin^2 \theta_W)$  and  $g_V^n = -\frac{1}{2}$  are the standard model weak constants. For a neutrino source that uses pion decay at rest, it contains  $\nu_\mu$ ,  $\bar{\nu}_\mu$ , and  $\nu_e$ . A coherent elastic scattering experiment can generate very small sensitivity to the  $\epsilon_{ee}^{qV}$  and  $\epsilon_{e\tau}^{qV}$  parameters that are currently allowed. [4]

The standard model also predicts that there is a neutrino magnetic moment,  $\mu_\nu \leq 10^{-19} \mu_B (m_\nu/1\text{eV})$ , which is very small compared to other limits that the standard model predicts. For example, the limit set from the lack of observed energy loss in red giant revolution's electromagnetic couplings is  $\mu_\nu \leq 10^{-12} \mu_B$ . [4] This small but non-zero magnetic moment can change the results of the coherent elastic scattering at low energies. The magnetic scattering cross section is given by[4]

$$\left( \frac{d\sigma}{dE} \right) = \frac{\pi \alpha^2 \mu_\nu^2 Z^2}{m_e^2} \left( \frac{1 - E/k}{E} + \frac{E}{4k^2} \right). \quad (6)$$

The change in the scattering events happen when neutrino magnetic moment is near the limit for  $\nu_e$  for recoil energies less than 10 keV, which may not be noticeable due to slightly higher thresholds for most detectors. For  $\nu_\mu$ , there could be a measurable difference if the detector has a 10 keV threshold. For nuclei with spin, they have



extra terms that could potentially change the cross sections and vary the scattering results.

The neutrino-nucleus coherent elastic scattering opens up possibilities for discovering new physics and improving some of the existing theories and models. Made possible by the invention of the low threshold detectors, the measurement of the scattering events has many applications, such as the detection of supernova neutrinos, reactor monitoring, measuring the Weinberg angle, and measuring form factors with precision which motivates this thesis.

### 1.3 Applications in Nuclear Physics and Astrophysics

A better understanding of neutron densities and neutron radii has many applications. One example is that obtaining a more accurate neutron density distribution can reduce uncertainties in atomic parity violation experiments.[1] Further, as mentioned previously, accurate measurements of neutron radii can improve the prediction of the energy functionals in different mean field theories significantly. Additionally, measuring neutron density distributions can help determine the neutron skin thickness, which has important implications in astrophysics.[2] Although the orbits and masses for neutron stars can be determined accurately, the radius of neutron stars and their moments of inertia can only be obtained through theories, which uses the equation of state for neutron-rich nuclear matter that relates to the density dependence of the nuclear symmetry energy. As mentioned in the first section, the symmetry energy is directly related to the neutron skin since neutron skins are sensitive to the symmetry energy at low densities and the neutron star radius is directly related to the symmetry energy at higher densities.[5] Therefore, measuring neutron density distributions provides information on the neutron skin and the equation of state of neutron matter, which helps determine the size of the neutron stars.

### 1.4 Detectors

There are many ongoing efforts around the world to conduct CE $\nu$ NS experiments. This section presents an overview of the detectors, their corresponding technologies, specifications, and anticipated energy thresholds that different laboratories worldwide implement for the CE $\nu$ NS experiments.

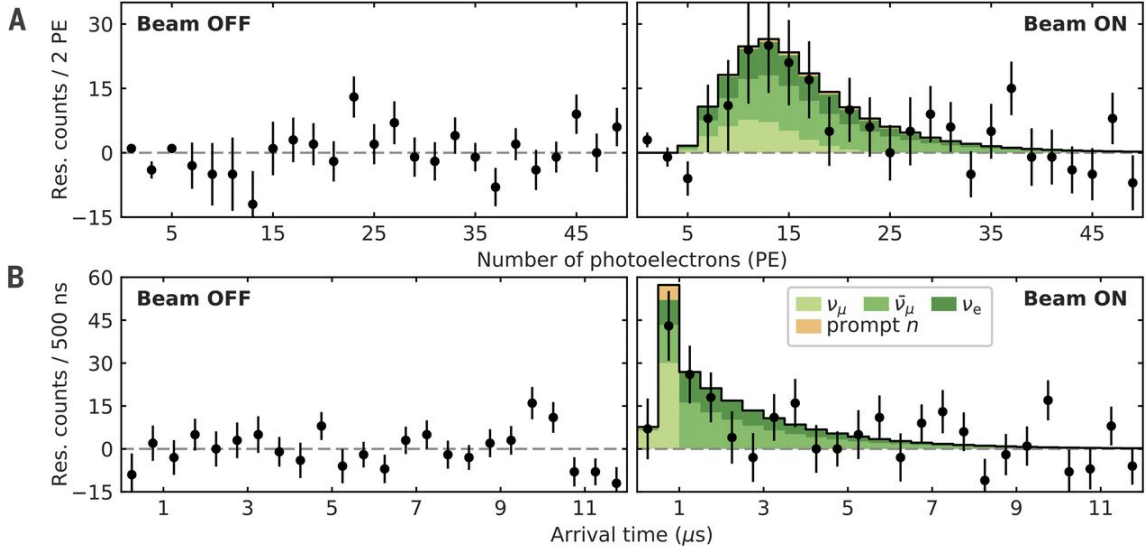


Figure 1: Observation of  $\text{CE}\nu\text{NS}$  events. Part A and B shows the residual differences between CsI signals 12 microseconds before and after the POT triggers. Approximately 1.17 photoelectrons are expected per keV of cesium or iodine nuclear recoil energy. Figure credit to Ref. [6]

The COHERENT experiment at the Spallation Neutron Source is the first and only experiment to successfully measure the  $\text{CE}\nu\text{NS}$  events. The results of the measurements are shown in Fig. 1. Approximately 1.17 photoelectrons are detected per keV of nuclear recoil energy.[6] The COHERENT experiment implements four different types of low-threshold/low-background technologies: CsI[Na] scintillating crystal, p-type point-contact germanium detectors, single-phase liquid argon, and NaI[Tl] scintillating crystals. The detailed parameters of the four subsystems is shown in Table 1. Fig. 2 shows the apparatus of the detector.[7]

Target	Technology	Mass (kg)	Distance from source (m)	Recoil threshold (keV)
CsI[Na]	Scintillating crystal	14.6	19.3	6.5
Ge	HPGe PPC	10	22	5
LAr	Single-phase	22	29	20
NaI[Tl]	Scintillating crystal	0.0925	28	13

Table 1: Parameters for the four subsystems of the COHERENT detector. Table credit to Ref. [7]

The NUCLEUS experiment at the Chooz nuclear power plant in France uses cryogenic detectors which are based on CRESST cryogenic detectors that have the world-leading sensitivity for searching low mass dark matter. The experiments have two phases: use a detector with 10g target for the first phase and use a detector with 1kg target for the second phase. The 10g target detector uses 6g of  $3 \times 3$  array of  $\text{CaWO}_4$  and 4g of  $3 \times 3$  array of  $\text{Al}_2\text{O}_3$ . Fig. 3 shows the detailed apparatus of the 10g target detector. The NUCLEUS detector has an unprecedentedly low energy threshold  $\leq 20\text{eV}$  and has the potential to have an eventual precision around 10% on the cross section of the coherent elastic scattering. The 1kg detector, planned to be ready for use in 2023, targets the measurement of the cross section with a 1% uncertainty. [8]

The CryoCube detector for RICOCHET aims to implement a compact array of 27 cubic 32g detectors using either germanium to build a semiconductor detector or use zinc to build a superconductor detector. The germanium detector allows a low energy threshold at around 10 eV while the zinc detector may achieve an extremely low threshold in the meV range. The detector also purposes to have a  $\geq 10^3$  EM background rejection and 1 kg target mass.[9]

This thesis is organized in the following way. Section II will talk about the methodology that we uses to simulate and calculate the coherent scattering events, including cross section, expansion of the form factor, and calculation of the effective moment. Section III will discuss the results generated from the methods in section II and perform a Monte-Carlo simulation to analyze the data in detail.

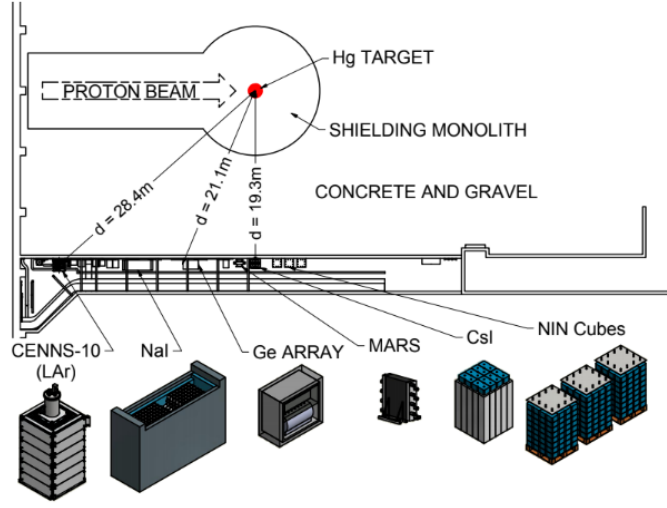


Figure 2: The setup of the four subsystems for the COHERENT detector. Figure credit to Ref. [7]

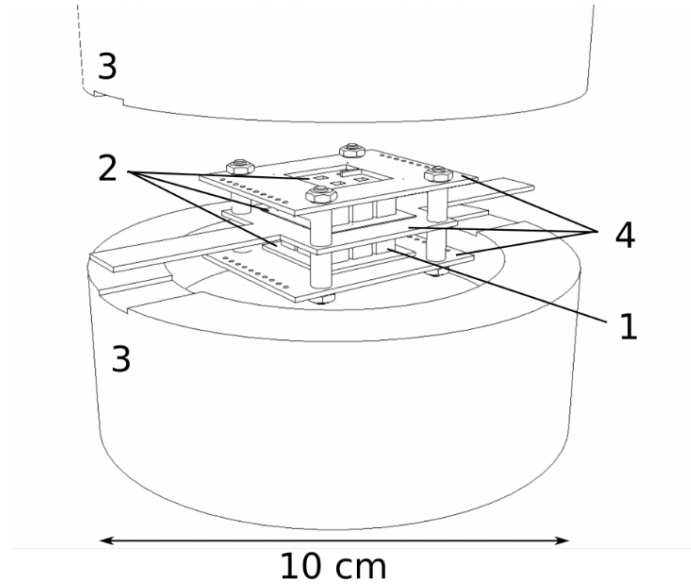


Figure 3: 3D sketch of the NUCLEUS-10g detector. it consists of three different kinds of cryogenic calorimeters - two  $3 \times 3$  arrays of gram-scale cryogenic calorimeters as the target (1), an inner veto (2) and an outer veto (3) with a 10cm diameter. A non-instrumented support structure (4) is implemented to hold the assembly. Figure credit to Ref. [8]

## 2 Methodology

In this section, we present the detailed steps to simulate the neutrino-nucleus coherent scattering, including the calculation of the cross section, the use and expansion of the form factor, and the calculation of effective moments for atoms like germanium that exist in isotopes with different abundances.

### 2.1 Calculation of the Form Factor and the Number of Scattering Events

For the neutrino-nucleus coherent elastic scattering, the cross section for a spherical and spin zero nucleus is given by [1]

$$\frac{d\sigma}{dT}(E, T) = \frac{G_F^2}{2\pi} \left[ 2 - \frac{2T}{E} + \left( \frac{T}{E} \right)^2 - \frac{MT}{E^2} \right] \frac{Q_W^2}{4} F^2(Q^2), \quad (7)$$

where  $G_F$  is the Fermi constant,  $E$  is the energy of the incoming neutrino,  $T$  is the recoil energy of the nucleus,  $M$  is the mass of the nucleus,  $Q_W = N - Z(1 - 4\sin^2\theta_W)$  is the weak charge of the nucleus ( $\sin^2\theta_W \approx 0.231$ ). The equation for the cross section contains the squared momentum transfer,  $Q^2 = 2E^2TM/(E^2 - ET)$ , and the form factor  $F(Q^2)$  as a function of  $Q^2$ , containing information about the nuclear densities by a Fourier transform.

The form factor for a spherical nucleus is given by [1]

$$F(Q^2) = \frac{1}{Q_W} \int [\rho_n(r) - (1 - 4\sin^2\theta_W)\rho_p(r)] \frac{\sin Q_r}{Q_r} r^2 dr, \quad (8)$$

where  $\rho_n$  and  $\rho_p$  are the neutron and proton densities respectively. The form factor modifies the non-coherent scattering at higher energies. Some effects such as the ones because of the finite size of the nucleon have been neglected, which changes the form factor at high moment transfer and modifies the neutron density. However, based on the sensitivity of our simulation, the differences that these effects cause are insignificant.

We can separate the neutron and proton parts in the form factor and obtain the following

$$F(Q^2) = \frac{1}{Q_W} \int [F_n(Q^2) - (1 - 4\sin^2\theta_W)F_p(Q^2)]. \quad (9)$$

Since  $1 - 4 \sin^2 \theta_W \approx 0.076$ , the contribution to the overall form factor from the proton term is quite small. Therefore, the scattering mainly depends on the neutrons and our simulations focus primarily on the neutron parts.

There are primarily two types of neutrino in our simulations, the ones from fission inside the nuclear reactors and the ones from decaying stopped pions. Since the reactor neutrinos have lower energies and the recoil energies generated by these neutrinos are hardly distinguishable from the background, we only look at the neutrinos from the pion decay. The pion decays through  $\pi^+ \rightarrow \nu_\mu + \mu^+$ . The muons can further decay and generate antineutrinos by  $\mu^+ \rightarrow e^+ + \nu_e + \bar{\nu}_\mu$ . We neglect the initial pion decay and only focus on the  $\bar{\nu}_\mu$  from  $\mu^+$ . The normalized spectra for the neutrinos  $\nu_e$  and antineutrinos  $\bar{\nu}_\mu$  are [2]

$$\begin{aligned} f_{\nu_e} &= \frac{96}{m_\mu^4} (m_\mu - E_{\nu_e}^2 - 2E_{\nu_e}^3) dE_{\nu_e}, \\ f_{\bar{\nu}_\mu} &= \frac{16}{m_\mu^4} (3m_\mu E_{\bar{\nu}_\mu}^2 - 4E_{\bar{\nu}_\mu}^3) dE_{\bar{\nu}_\mu}, \end{aligned} \tag{10}$$

where  $m_\mu$  is the muon mass. These two equations give the probability density of finding an emitted neutrino between energy  $E$  and  $E + dE$ . The energy spectrum of the two fluxes is shown below in Fig. 4.

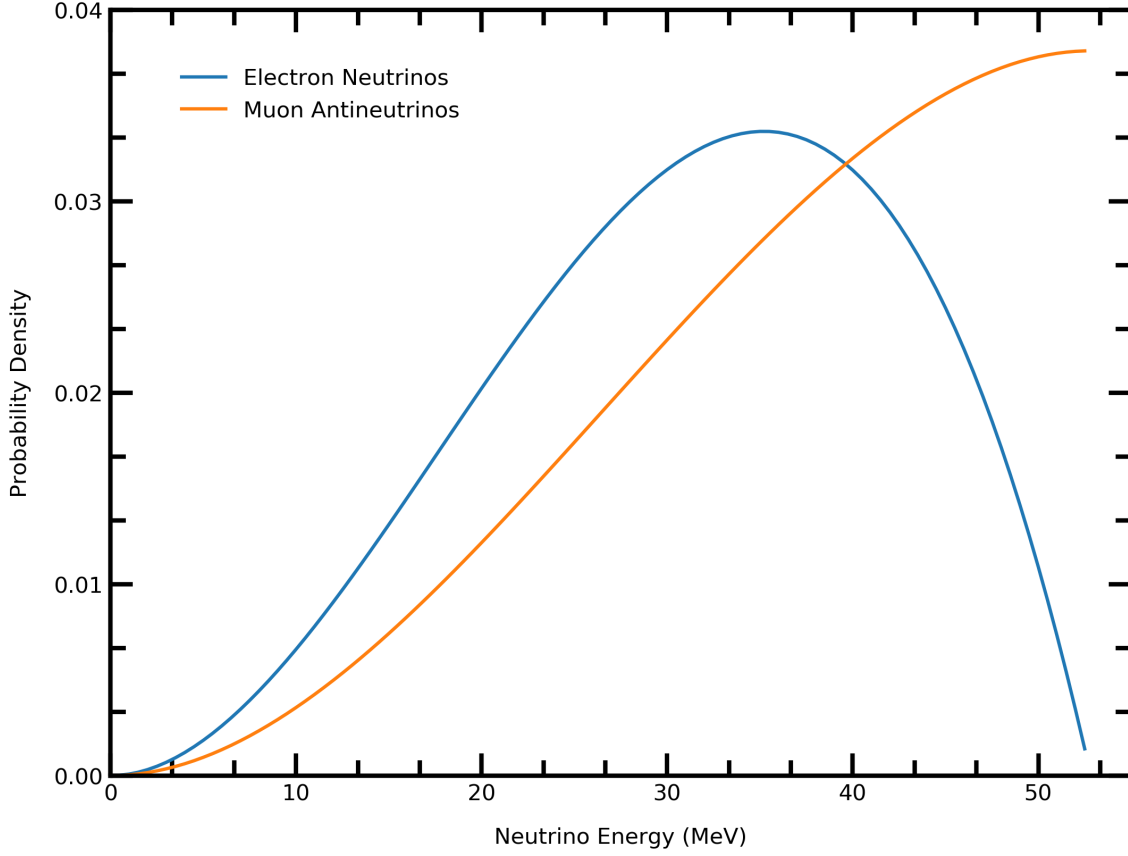


Figure 4: The probability density of finding electron neutrinos and muon antineutrinos at different neutrino energies.

Combining the cross section with the neutrino flux, we can calculate the number of scattering events based on the recoil energy: [2]

$$\frac{dN}{dT}(T) = N_t C \int_{E_{min}(T)}^{m_\mu/2} f(E) \frac{d\sigma}{dT}(E, T) dE, \quad (11)$$

where  $N_t$  is the number of targets in the detector,  $C$  is the total number of neutrinos of a specific flavor approaching the target per second per cm squared,  $E_{min}(T) = \frac{1}{2}(T + \sqrt{T^2 + 2TM})$  is the minimum energy required for a neutrino to generate the recoil energy  $T$  at the nucleus,  $m_\mu/2$  is the maximum energy from a neutrino, and  $f(E)$  is the sum of the fluxes for all types of neutrinos produced.



## 2.2 Form Factor

Using the analytic method from Ref.[10], we can represent form factors using the analytic approximations of the density distributions instead of nuclear structure calculations, which simulates the nucleons with a thickness  $s$  on the surface and a constant interior density. The skin thickness  $s$  represents how fast the density falls off from the interior to zero. We obtain the form factors by the Fourier transform of the density distributions, also called the Helm form factor: [10]

$$F(Q^2) = \frac{3j_1(QR_0)}{QR_0} \exp\left[-\frac{1}{2}(Qs)^2\right], \quad (12)$$

where  $j_1$  represents the first spherical Bessel function,  $R_0^2 = R^2 - 5s^2$  is around 3 fm to 5 fm,  $R$  is the radius of the nucleus, and  $s$  is the thickness of the surface. Fig. 5 shows the relationship between the Helm form factor and the momentum transfer  $Q$  with different possible  $R_0$ . Using Eq. (12) to calculate the form factors for protons and neutrons, we can get the full form factor by incorporating the two parts into Eq. (9).

For detectors filled with  $^{40}\text{Ar}$  where  $N = 22$  and  $Z = 18$ , we use the data from Ref.[11], which gives the value for the mean square radius of the proton,  $\langle R_p^2 \approx 11.75 \text{ fm}^2 \rangle$ . We then take the square root of the value and get the proton radius,  $R_p = 3.43 \text{ fm}$ . For the surface thickness, we set the same value,  $s = 0.5 \text{ fm}$ . [1] Since the purpose of our simulation is to demonstrate the possibility of conducting such an experiment in reality, it is adequate to assign the values to the parameters to generate an analytic model for demonstration while specific calculations and values need to be made and determined for performing a real experiment. For a nucleus that has the same number of protons and neutrons, the proton and neutron density distribution should be roughly the same, which means that  $R_p \approx R_n$ . Therefore, for approximations, we can set  $R_p = R_n$ , which means that we can compute both  $F_n(Q^2)$  and  $F_p(Q^2)$  using Eq. (12). The calculated Helm form factor for argon is shown in Fig. 6. We can then use Eq. (9) to calculate the full form factor, which we can plug into Eq. (7) to compute the cross section. With the cross section calculated, we can combine it with the neutrinos flux from Eq. (10) to simulate the scattering events by integrating over Eq. (11) and plot it out as a base curve for reference, which is demonstrated in Fig. 7.

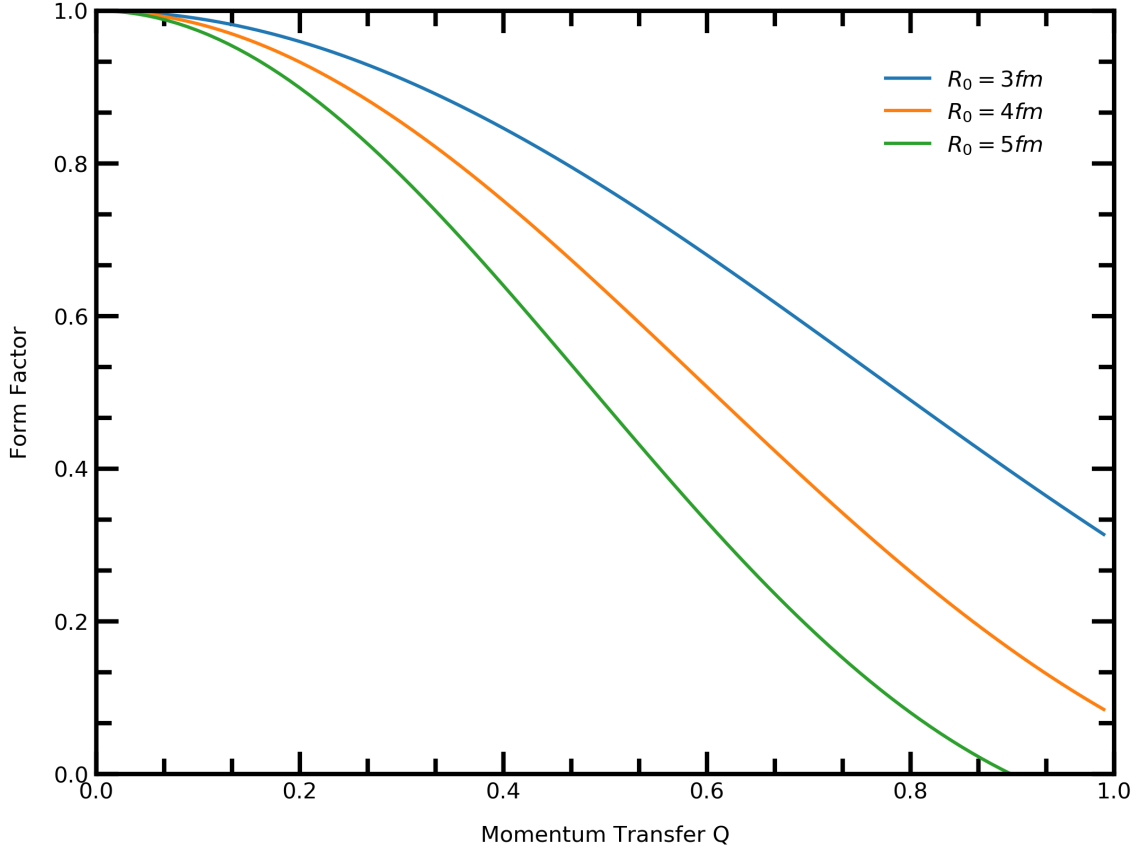


Figure 5: Helm form factor against momentum transfer  $Q$  at  $R_0 = 3$  fm,  $R_0 = 4$  fm, and  $R_0 = 5$  fm. The variation in the  $R_0$  value changes the shape of the curve for the form factor.

### 2.3 Form Factor Expansion

Since  $Q$  is very small, we can apply Taylor expansion to the form factor and truncate the series so that we can determine the form factor using only a few parameters that we can obtain from the nuclear structure calculations. The Taylor expansion of the neutron form factor is [2]

$$\begin{aligned}
 F_n(Q^2) &\approx \int \rho_n(r) \left( 1 - \frac{Q^2}{3!} r^2 + \frac{Q^4}{5!} r^4 - \frac{Q^6}{7!} r^6 + \dots \right) r^2 dr \\
 &\approx N \left( 1 - \frac{Q^2}{3!} \langle R_n^2 \rangle + \frac{Q^4}{5!} \langle R_n^4 \rangle - \frac{Q^6}{7!} \langle R_n^6 \rangle + \dots \right)
 \end{aligned} \tag{13}$$

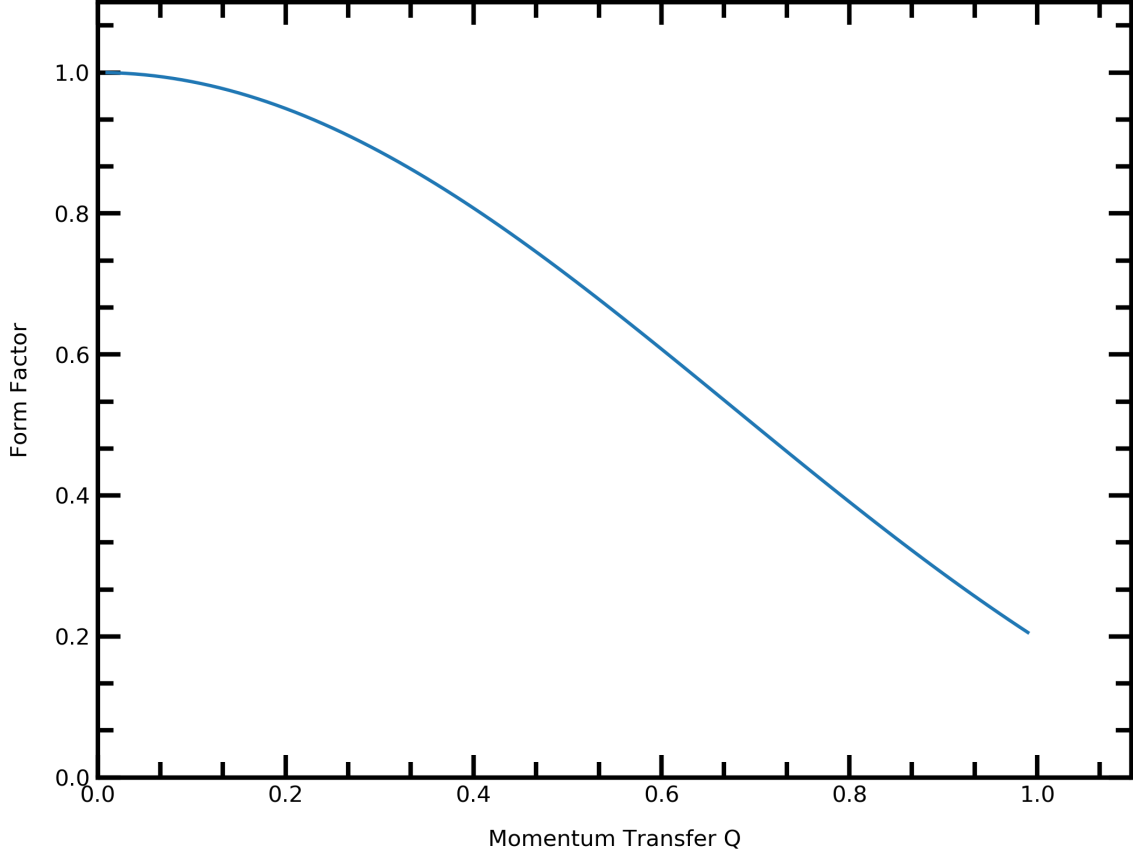


Figure 6: Helm form factor against momentum transfer  $Q$  for  $^{40}\text{Ar}$ . We set  $R_0 = 3.43$  for both proton and neutron full factors,  $N = 22$ ,  $Z = 18$ , and  $s = 0.5$  fm.

where

$$\langle R_n^k \rangle = \frac{\int \rho_n r^k d^3r}{\int \rho_n d^3r}. \quad (14)$$

Here we only focus on the neutron part of the form factor since the proton terms will barely change the result. After the expansion, the form factor consists only of the even moments of the neutron density distributions, which are easy to calculate and represent physical and measurable parameters.<sup>[2]</sup> Due to the low energy of the neutrinos, we can omit the terms after the first two terms for nuclei with lighter weights such as Ar and Ge and after the first three terms for heavier ones like Xe. Fig. 8 shows the comparison between the results obtained from the Helm form factor and the truncated form factor from Taylor expansion. The difference in the counts is noticeable but gradually becomes smaller at lower energies from 0 to 60 keV, and approaches 0 as the energy goes higher than 60 keV, suggesting that the truncated

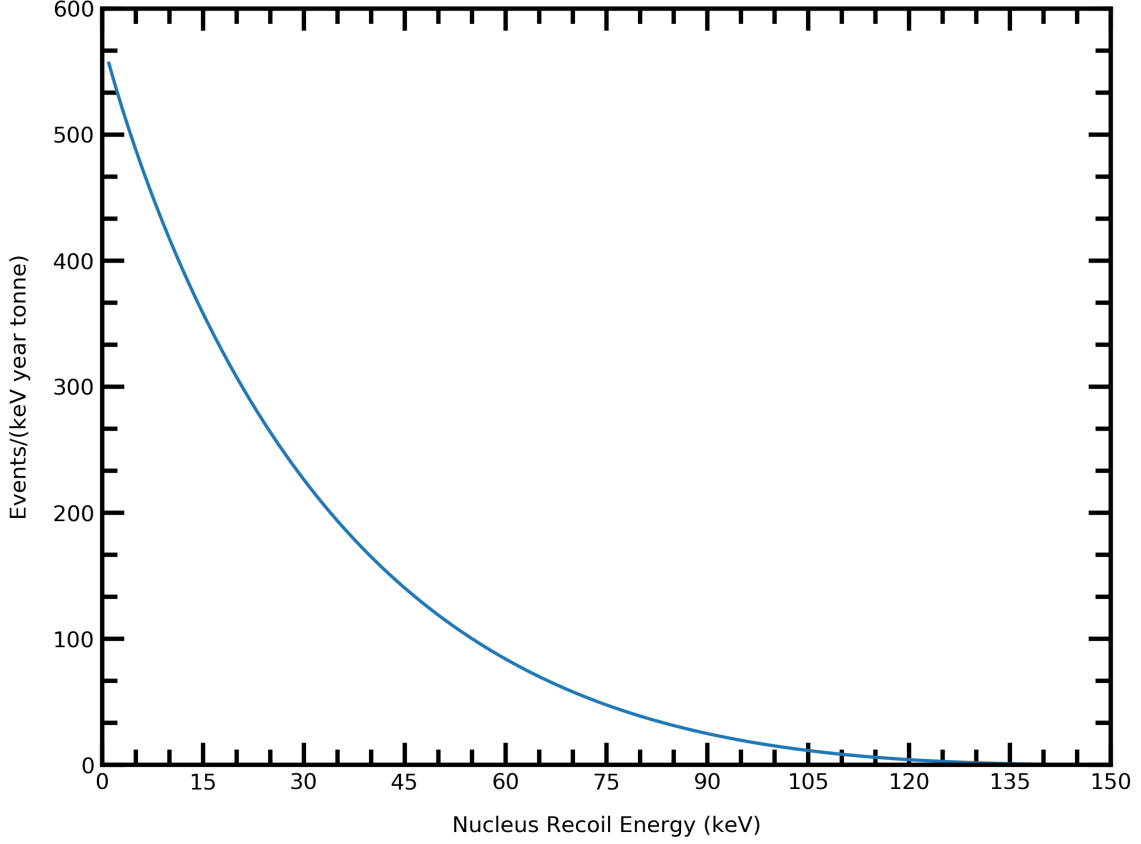


Figure 7: Number of  $\nu_e$  and  $\bar{\nu}_\mu$  events in detector filled with  $^{40}\text{Ar}$  in units of number per (keV year tonne) assuming  $10^7$  neutrinos of the two flavors are emitted from the sources every second. Assume  $R_n = R_p$ .

series is an applicable implementation that preserves the accuracy of the result. Fig. 9 shows the results at different  $R$  values. The changes in  $R$  have enough impacts on the results to the curves distinguishable.

## 2.4 Effective Moments

When previously generating the base curve for Ar, we used  $^{40}\text{Ar}$  to calculate the number of scattering event because 99.6% of Ar is constituted of  $^{40}\text{Ar}$ . However, for elements such as germanium and xenon, they exist in multiple isotope forms with sufficient abundances that we need to define the effective moments in order to include the different isotopes. The isotopes and abundances for germanium and xenon are shown in Table 2.

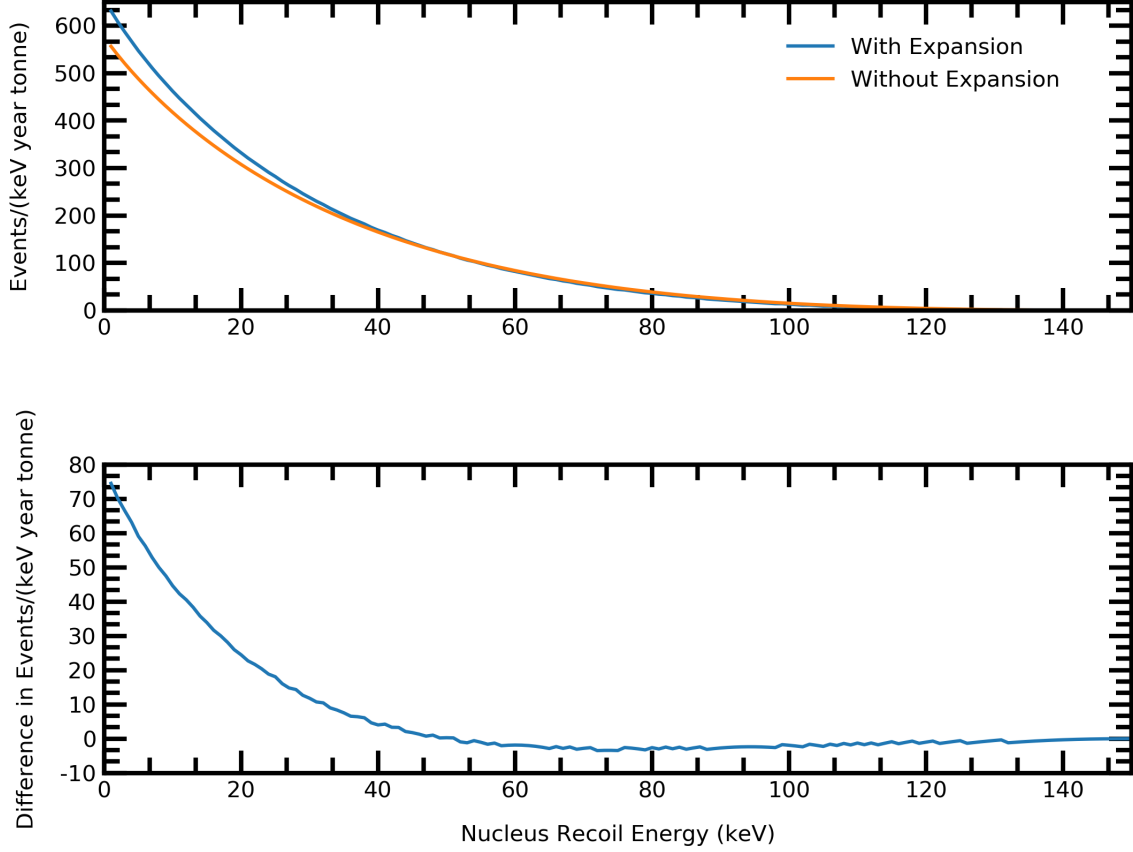


Figure 8: Number of  $\nu_e$  and  $\bar{\nu}_\mu$  events in detector in units of number per (keV year tonne) using Helm form factor and the first two terms of Taylor expansion of the form factor neglecting the proton part. The subplot below shows the difference in counts between the two methods.

To account for the isotopes, we need to sum over all the isotopes. The equation for calculating the scattering events then becomes: [2]

$$\frac{dN}{dT}(T) = N_A M_{\text{detector}} C \int f(E) \sum_i \left[ \frac{X_i}{M_i} \left( \frac{d\sigma}{dT}(T, E) \right)_i \right] dE, \quad (15)$$

where  $N_A M_{\text{detector}}$  substitutes the  $N_t$  in Eq. 11 and the summation adds up all the isotopes.  $X_i$  is the natural abundance of isotope  $i$ ,  $M_i$  is the mass of the isotope,  $N_A$  is the Avogadro's number, and  $M_{\text{detector}}$  is the total mass of the element in the detector.[2]

For simulating the scattering event for germanium, neglecting the proton terms and focusing only on the neutron terms, we can use the first two terms of the expansion

Eq. 13 and plug them into the cross section given by Eq. 7 for calculating the sum of the cross sections for all isotopes in Eq. 15. We then get:

$$\begin{aligned}
\sum_i \left[ \frac{X_i}{M_i} \left( \frac{d\sigma}{dT}(T, E) \right)_i \right] = & \frac{G_F^2}{8\pi} \left[ \sum_i (X_i N_i^2) \left( 2 - \frac{2T}{E} + \left( \frac{T}{E} \right)^2 \right) - \sum_i (X_i N_i^2 M_i) \left( \frac{T}{E^2} \right) \right. \\
& - \sum_i (X_i N_i^2 M_i \langle R_n^2 \rangle_i) \left( 2 - \frac{2T}{E} + \left( \frac{T}{E} \right)^2 \right) \frac{Q^2}{3\langle M \rangle} \\
& + \sum_i (X_i N_i^2 M_i^2 \langle R_n^2 \rangle_i) \left( \frac{T}{E^2} \right) \frac{Q^2}{3\langle M \rangle} \\
& + \sum_i (X_i N_i^2 M_i^2 \langle R_n^2 \rangle_i^2) \left( 2 - \frac{2T}{E} + \left( \frac{T}{E} \right)^2 \right) \frac{Q^4}{36\langle M \rangle^2} \\
& - \sum_i (X_i N_i^2 M_i^3 \langle R_n^2 \rangle_i^2) \left( \frac{T}{E^2} \right) \frac{Q^4}{36\langle M \rangle^2} \\
& + \sum_i (X_i N_i^2 M_i^2 \langle R_n^4 \rangle_i) \left( 2 - \frac{2T}{E} + \left( \frac{T}{E} \right)^2 \right) \frac{Q^4}{60\langle M \rangle^2} \\
& - \sum_i (X_i N_i^2 M_i^3 \langle R_n^4 \rangle_i) \left( \frac{T}{E^2} \right) \frac{Q^4}{60\langle M \rangle^2} \\
& - \sum_i (X_i N_i^2 M_i^3 \langle R_n^2 \rangle_i \langle R_n^4 \rangle_i) \left( 2 - \frac{2T}{E} + \left( \frac{T}{E} \right)^2 \right) \frac{Q^6}{360\langle M \rangle^3} \\
& \left. + \sum_i (X_i N_i^2 M_i^4 \langle R_n^2 \rangle_i \langle R_n^4 \rangle_i) \left( \frac{T}{E^2} \right) \frac{Q^6}{360\langle M \rangle^3} + \dots \right], \tag{16}
\end{aligned}$$

where  $\langle M \rangle = \sum_i X_i M_i$  and  $Q^2 = 2E^2 T \langle M \rangle / (E^2 - ET)$ . Here we are keeping the

Isotope	Abundance	Isotope	Abundance
$^{70}\text{Ge}$	0.205	$^{128}\text{Xe}$	0.0191
$^{72}\text{Ge}$	0.274	$^{129}\text{Xe}$	0.264
$^{73}\text{Ge}$	0.078	$^{130}\text{Xe}$	0.041
$^{74}\text{Ge}$	0.365	$^{131}\text{Xe}$	0.212
$^{76}\text{Ge}$	0.205	$^{132}\text{Xe}$	0.269
		$^{134}\text{Xe}$	0.104
		$^{136}\text{Xe}$	0.089

Table 2: Isotopes and abundances for germanium and xenon[12]

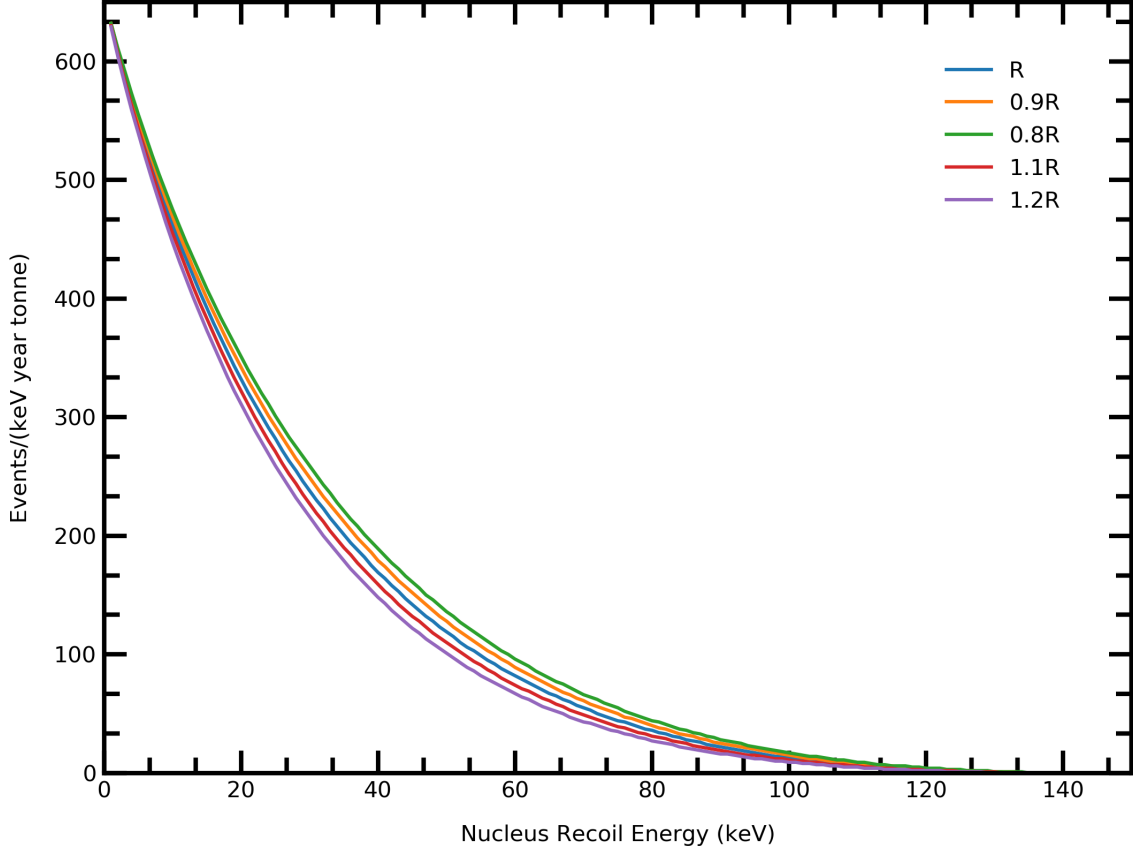


Figure 9: Number of  $\nu_e$  and  $\bar{\nu}_\mu$  events in detector in units of number per (keV year tonne) for  $^{40}\text{Ar}$  at slightly different  $R$  values. The base values are  $\langle R_n^2 \rangle^{1/2} = 3.417$  fm and  $\langle R_n^4 \rangle^{1/4} = 3.723$  fm, adopted from the SkM\* standard values.[14] The 10% increments and decrements only slightly vary the resulting curves.

$\langle R_n^2 \rangle$  and  $\langle R_n^4 \rangle$  terms for germanium since we truncate the series for the neutron form factor after the first two terms that are consistent with the  $\langle R_n^2 \rangle$  and  $\langle R_n^4 \rangle$  terms. For heavier elements like xenon, terms up to  $\langle R_n^6 \rangle$  need to be preserved. The two effective second moments are [2]

$$\begin{aligned} \langle R_n^2 \rangle_{eff,1} &= \frac{\sum_i X_i N_i^2 M_i \langle R_n^2 \rangle_i}{\sum_i X_i N_i^2 M_i}, \\ \langle R_n^2 \rangle_{eff,2} &= \frac{\sum_i X_i N_i^2 M_i^2 \langle R_n^2 \rangle_i}{\sum_i X_i N_i^2 M_i^2}, \end{aligned} \tag{17}$$

which are derived from the third and fourth terms in Eq. 16, and the two effective

fourth moments are [2]

$$\begin{aligned}\langle R_n^4 \rangle_{eff,1} &= \frac{\sum_i X_i N_i^2 M_i^2 \langle R_n^4 \rangle_i}{\sum_i X_i N_i^2 M_i^2}, \\ \langle R_n^4 \rangle_{eff,2} &= \frac{\sum_i X_i N_i^2 M_i^3 \langle R_n^4 \rangle_i}{\sum_i X_i N_i^2 M_i^3},\end{aligned}\tag{18}$$

which are derived from the seventh and eighth terms in Eq. 16.

Since the differences between the two effective second moments and the two effective fourth moments are negligibly small, we can assume that the two equations give the same result. Plugging the effective moments equations into Eq. 16, we get the new form for the the cross section, which only differs the original form by 0.01%, [2]

$$\begin{aligned}\sum_i \left[ \frac{X_i}{M_i} \left( \frac{d\sigma}{dT}(T, E) \right)_i \right] &= \frac{G_F^2}{8\pi} \left[ \sum_i (X_i N_i^2) \left( 2 - \frac{2T}{E} + \left( \frac{T}{E} \right)^2 \right) - \sum_i (X_i N_i^2 M_i) \left( \frac{T}{E^2} \right) \right. \\ &\quad - \langle R_n^2 \rangle_{eff} \sum_i (X_i N_i^2 M_i) \left( 2 - \frac{2T}{E} + \left( \frac{T}{E} \right)^2 \right) \frac{Q^2}{3\langle M \rangle} \\ &\quad + \langle R_n^2 \rangle_{eff} \sum_i (X_i N_i^2 M_i^2) \left( \frac{T}{E^2} \right) \frac{Q^2}{3\langle M \rangle} \\ &\quad + \langle R_n^2 \rangle_{eff}^2 \sum_i (X_i N_i^2 M_i^2) \left( 2 - \frac{2T}{E} + \left( \frac{T}{E} \right)^2 \right) \frac{Q^4}{36\langle M \rangle^2} \\ &\quad - \langle R_n^2 \rangle_{eff}^2 \sum_i (X_i N_i^2 M_i^3) \left( \frac{T}{E^2} \right) \frac{Q^4}{36\langle M \rangle^2} \\ &\quad + \langle R_n^4 \rangle_{eff} \sum_i (X_i N_i^2 M_i^2) \left( 2 - \frac{2T}{E} + \left( \frac{T}{E} \right)^2 \right) \frac{Q^4}{60\langle M \rangle^2} \\ &\quad - \langle R_n^4 \rangle_{eff} \sum_i (X_i N_i^2 M_i^3) \left( \frac{T}{E^2} \right) \frac{Q^4}{60\langle M \rangle^2} \\ &\quad - \langle R_n^2 \rangle_{eff} \langle R_n^4 \rangle_{eff} \sum_i (X_i N_i^2 M_i^3) \left( 2 - \frac{2T}{E} + \left( \frac{T}{E} \right)^2 \right) \frac{Q^6}{360\langle M \rangle^3} \\ &\quad \left. + \langle R_n^2 \rangle_{eff} \langle R_n^4 \rangle_{eff} \sum_i (X_i N_i^2 M_i^4) \left( \frac{T}{E^2} \right) \frac{Q^6}{360\langle M \rangle^3} + \dots \right].\end{aligned}\tag{19}$$

Adopting the new expression of the cross section that sums over all isotopes from



Eq. 19, we can simulate and make predictions on the number of scattering events for detectors filled with germanium using the data shown in Table 3, as shown in Fig. 10.

Remember that when deriving the cross section for light elements with multiple isotopes, we neglect the proton terms since it has very small impact on our results. In order to make the result more accurate, we can add back the proton terms and the terms including both protons and neutrons. Remember that we previously separate the two parts of the full form factor using Eq. 9, which gives us

$$F^2(Q^2) = F_n^2(Q^2) - 2\alpha F_n(Q^2)F_p(Q^2) + \alpha^2 F_p^2(Q^2) \quad (20)$$

where  $\alpha = 1 - 4\sin^2\theta_W$ . Since the  $F_p^2(Q^2)$  terms are the same as the  $F_n^2(Q^2)$  terms shown in Eq. 19 with all the parameters relevant to neutrons changing to the ones with protons in them, we only need simple substitution of some constants to obtain the  $F_p^2(Q^2)$  part. For the part with both protons and neutrons, we only need to add the cross term,

$$\begin{aligned} F_n(Q^2)F_p(Q^2) = & 1 - \frac{Q^2}{3!} (\langle R_n^2 \rangle + \langle R_p^2 \rangle) + \frac{Q^4}{5!} (\langle R_n^4 \rangle + \langle R_p^4 \rangle) + \frac{Q^4}{(3!)^2} (\langle R_n^2 \rangle \langle R_p^2 \rangle) \\ & - \frac{Q^6}{3! \cdot 5!} (\langle R_n^2 \rangle \langle R_p^4 \rangle + \langle R_n^4 \rangle \langle R_p^2 \rangle) + \frac{Q^8}{(5!)^2} (\langle R_n^4 \rangle \langle R_p^4 \rangle), \end{aligned} \quad (21)$$

which is obtained by multiplying the terms up to  $Q^4$  in the neutron form factor with the ones from the proton form factor. The coefficient

$$\frac{N_i^2 X_i M_i^k}{\langle M \rangle^k} \quad (22)$$

then becomes

$$\frac{N_i Z_i X_i M_i^k}{\langle M \rangle^k} \quad (23)$$

for the cross term. Adding the proton term and the cross term, we can get the cross section with the full form factor. We can then simulate the scattering event using the new cross section and get a more accurate curve that contains slightly fewer events at each energy bin compared to the previous one. Fig. 10 shows the result and compares it with the one that we previously obtained using only the neutron form factor. The

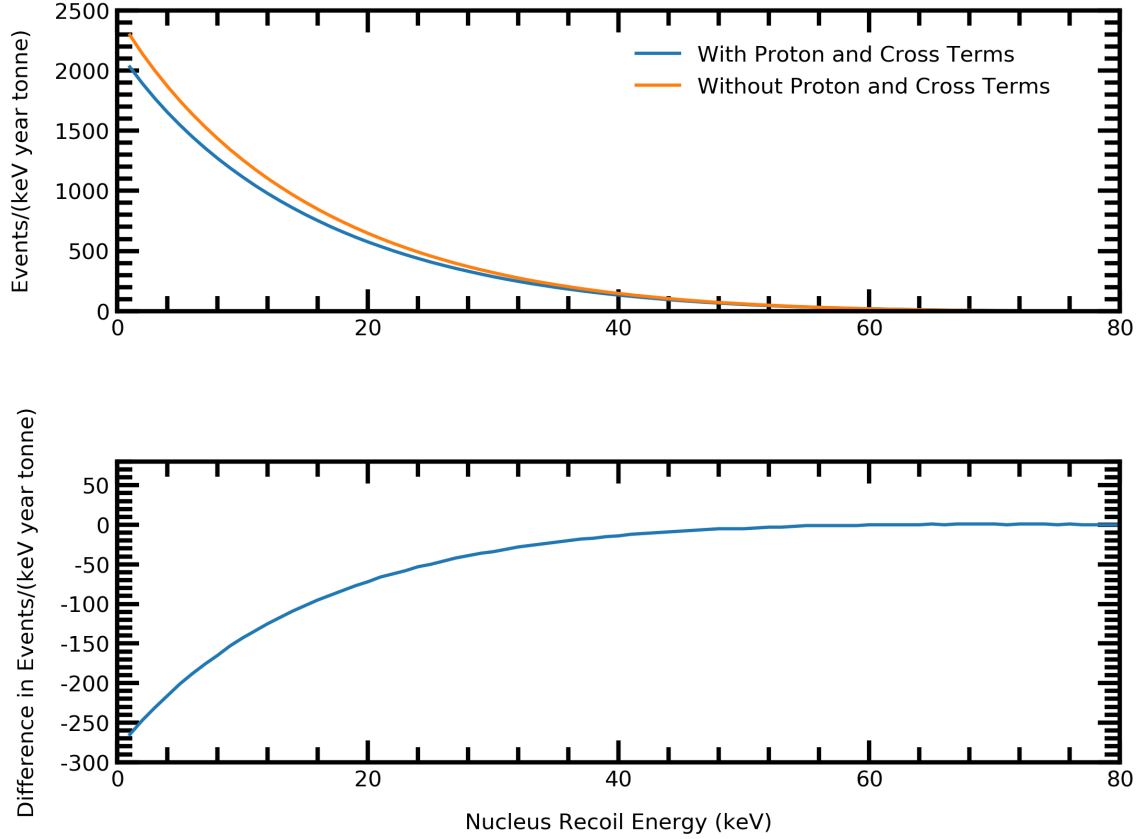


Figure 10: Number of  $\nu_e$  and  $\bar{\nu}_\mu$  events in detector in units of number per (keV year tonne) for *Ge* using only the neutron form factor and for adding the proton and cross terms. The subplot below shows the difference in results between the two methods.

difference is quite large at lower energies, but goes to zero as the energy increases to around 50 keV.

In order to predict different neutron distributions for different nuclear structure calculations and predict the number of scattering events expected in the detector, we need to vary the neutron radius. We vary the radius from  $-20\%$  to  $+20\%$  with an incremental step of  $10\%$  and create a database containing the scattering events at each radius. We then run a Monte-Carlo simulation for our data as an analysis, which we will discuss in the next section.

Atomic Mass (amu)	N	Abundance	$\langle R_p^2 \rangle$ (fm <sup>2</sup> )	$\langle R_p^4 \rangle$ (fm <sup>4</sup> )	$\langle R_n^2 \rangle$ (fm <sup>2</sup> )	$\langle R_n^4 \rangle$ (fm <sup>4</sup> )
69.924	38	0.206	15.428	323.689	15.832	343.030
71.922	40	0.275	15.895	344.108	16.568	377.481
72.923	41	0.078	15.665	330.939	16.457	368.242
73.921	42	0.365	15.974	345.718	16.893	390.827
75.921	44	0.077	16.026	345.900	17.170	401.341

Table 3: Atomic mass, number of neutrons, abundance,  $\langle R_p^2 \rangle$ ,  $\langle R_p^4 \rangle$ ,  $\langle R_n^2 \rangle$ , and  $\langle R_n^4 \rangle$  values for germanium isotopes. Table obtained from Ref. [13].

## 3 Results and Discussion

### 3.1 Binned Scattering Events

Instead of showing the counts at every keV, we present and store our results in a more realistic way. We add up the counts in every 5 keV range together and represent our scattering curves as binned bar plots. This is because most detectors do not have resolutions small enough to distinguish the difference in 1 keV. Most simulations similar to ours, such as the ones in Ref. [1] and Ref. [2], use 10 keV as the energy bin and resolution. Therefore, our implementation of the 5 keV energy bin helps determine whether a finer resolution can lead to a more satisfactory result. The detailed method using Python is shown in Appendix A. Examples of our binned events are shown in Fig. 11, Fig. 12, and Fig. 13. Comparing Fig. 11 and Fig. 12, we notice that *Ar* goes to a higher recoil energy than *Ge* since *Ar* has a lighter atomic mass. *Ge* has more counts overall because the cross section is proportional to  $N^2$  and *Ge* contains more neutrons in the atom. In Fig. 12 and Fig. 13, we can again see the effect of the added proton and cross terms in the form factor since the scattering events at each bin decrease.

### 3.2 Monte-Carlo Simulations

With the simulations running correctly, we then perform a simple Monte Carlo simulation to test the accuracy of determining the nuclear moments via the scattering events. We assume that the detector is filled with 10 tonnes of natural germanium and experiences a flux from pion decay of  $3 \times 10^7$  neutrinos per second per cm squared in each flavor for one year. According to Ref. [2], the neutrino production rate at multiple laboratories such as the Spallation Neutron Source at Oak Ridge National Laboratory, DAE $\delta$ ALUS, and the European Spallation Source varies from  $1 \times 10^{15}$  neutrinos per second per  $cm^2$  to  $3.5 \times 10^{15}$  neutrinos per second per  $cm^2$ . Therefore, our simulation represents placing the detectors from 16 to 30 meters from the neutrino sources, depending on the setup at different laboratories.[2]

To set up our Monte Carlo simulation, we first create a database of the scattering events from 5 to 80 keV and binned with a 5 keV bin size. The lower cutoff at 5 keV is due to the detector's possible threshold. To generate a database with sufficient data, we vary the  $\langle R_n^2 \rangle$  value from  $-15\%$  to  $+15\%$  and vary the  $\langle R_n^4 \rangle$  value from  $-30\%$  to

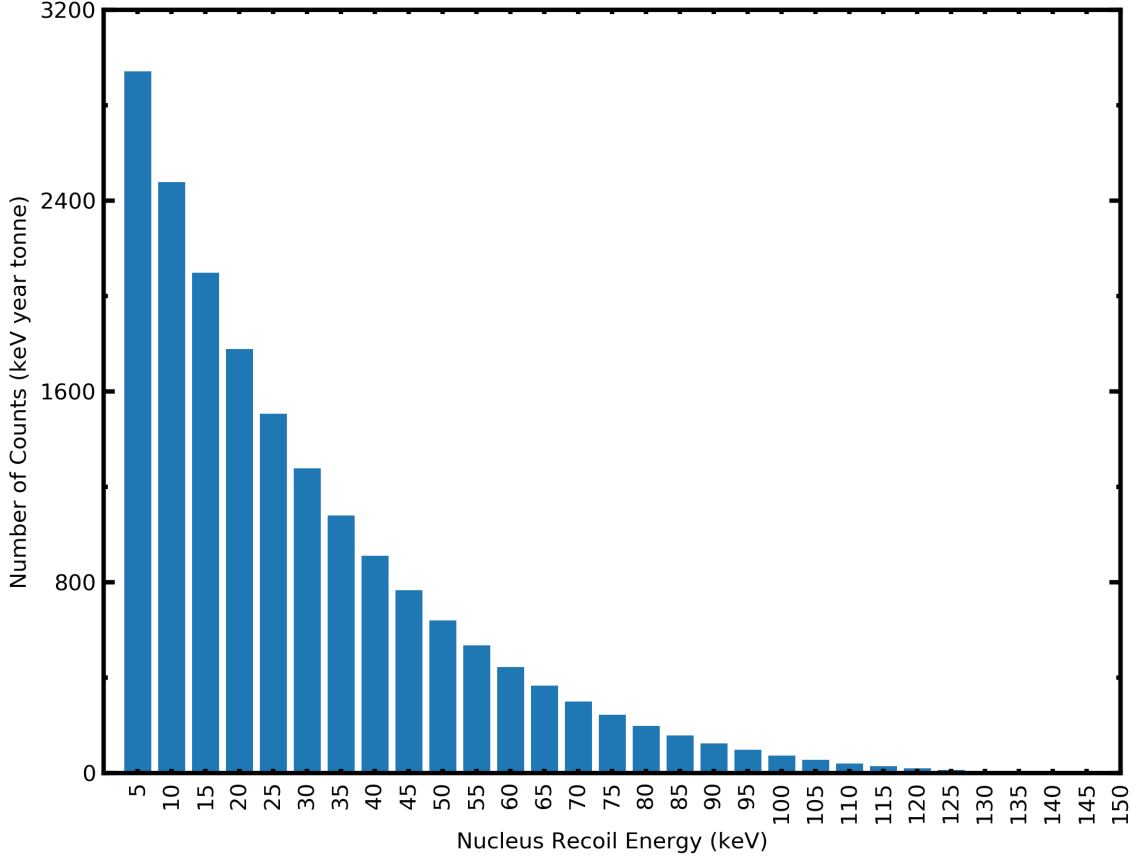


Figure 11: Number of  $\nu_e$  and  $\bar{\nu}_\mu$  events per (keV year tonne) summed in 5 keV bins for  $^{40}\text{Ar}$ .

+30%. Each value changes in a small step to give 100 different values. Combining the  $\langle R_n^2 \rangle$  and the  $\langle R_n^4 \rangle$  values, our database contains the scattering events for 10000 scenarios of different  $\langle R_n^2 \rangle$  and  $\langle R_n^4 \rangle$  values. Our Monte Carlo simulation assumes that the detection efficiency is always 100% and the neutrino flux is held constant so that we do not need to change the flux parameter during our simulation. For each run, we take the standard scattering events computed using the SkM\*  $\langle R_n^2 \rangle$  and the  $\langle R_n^4 \rangle$  values and add random noise to each bin of the scattering events, which we present mathematically as [14]

$$N'_{events} = N_{events} + \sqrt{N_{events}}s, \quad (24)$$

where  $N_{events}$  is the standard scattering events obtained using the SkM\* values[14] and  $s$  is a random number between -1 and 1 generated from a Gaussian distribution.

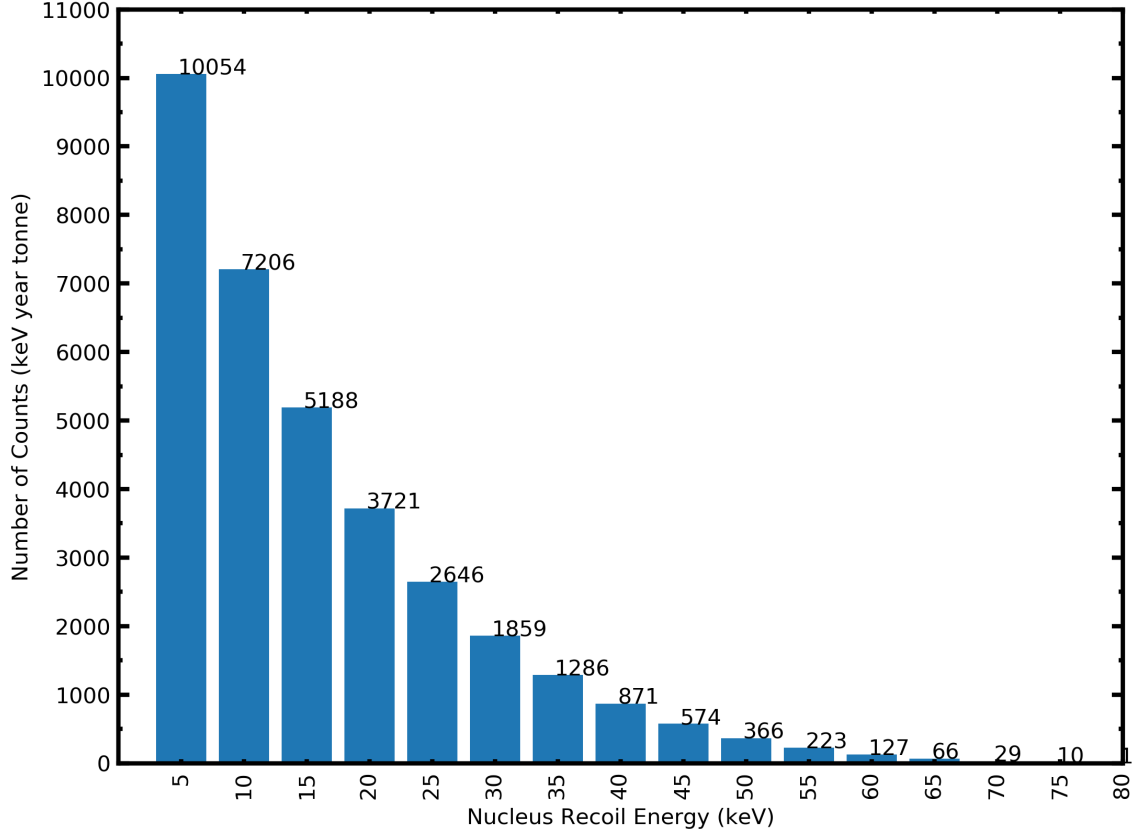


Figure 12: Number of  $\nu_e$  and  $\bar{\nu}_\mu$  events per (keV year tonne) summed in 5 keV bins for  $Ge$ , calculated using only the neutron terms in the form factor.

We then take the events with random noise and calculate the  $\chi^2$  value for each of 10000 events in our database using

$$\chi^2 = \sum \frac{(N'_{events} - N_{database})^2}{N_{database}}. \quad (25)$$

We then find the corresponding  $\langle R_n^2 \rangle$  and the  $\langle R_n^4 \rangle$  pair in our database that give the smallest  $\chi^2$  value. We repeat this process 5000 times and analyze the trends in the results. Fig. 14 shows the scatter plot of the 5000  $\langle R_n^2 \rangle$  and the  $\langle R_n^4 \rangle$  pairs.

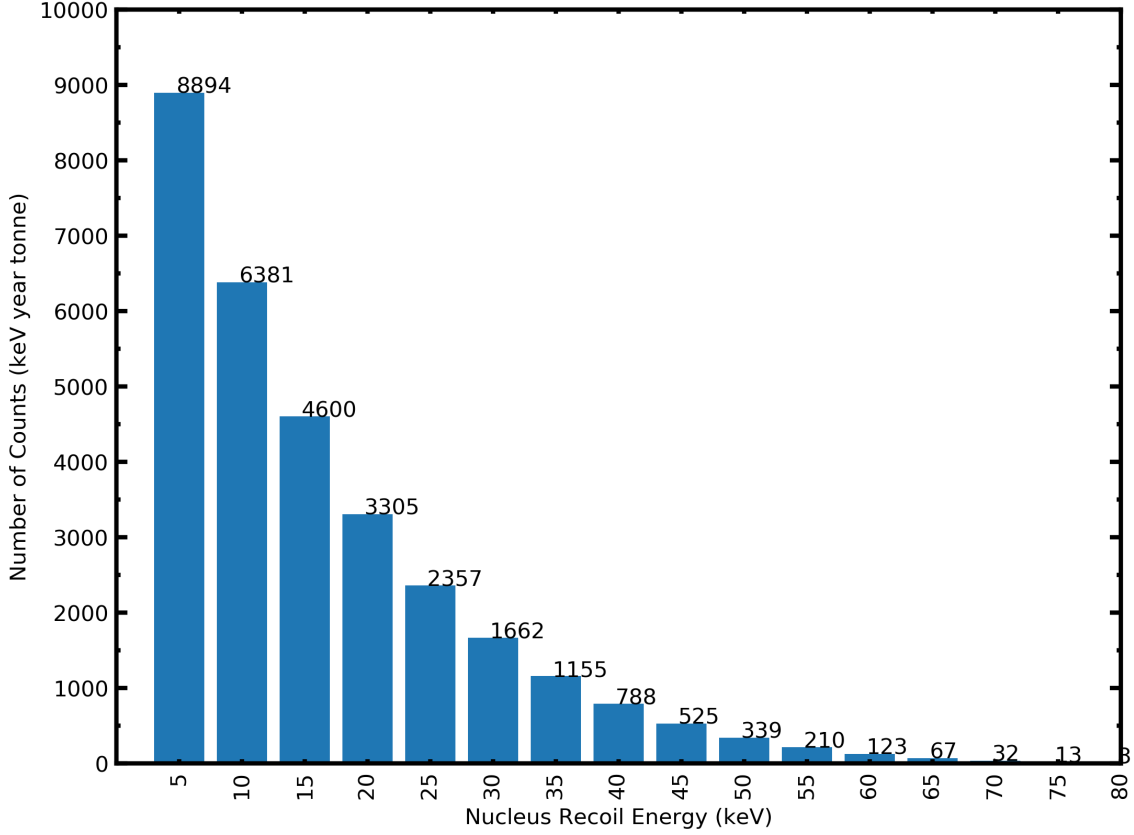


Figure 13: Number of  $\nu_e$  and  $\bar{\nu}_\mu$  events per (keV year tonne) summed in 5 keV bins for  $Ge$ , calculated using the full form factor with the proton and cross terms.

### 3.3 Discussion

From Fig. 14, we can see that the range of possible  $\langle R_n^4 \rangle^{1/4}$  values is a lot larger than that of  $\langle R_n^2 \rangle^{1/2}$  values. This issue about the dependence on  $\langle R_n^4 \rangle^{1/4}$  can be explained by looking at the contribution to the recoil distributions of the effective fourth moment in the form factors. For germanium, a 10% change in the fourth moment will lead to 1.3% more scattering events while a 10% change in  $\langle R_n^2 \rangle^{1/2}$  will result in a 6% change in scattering events.[2] Since the effective second moment relates to the nuclear recoil energy more closely, its value varies in a smaller range and has a lower uncertainty compared to the  $\langle R_n^2 \rangle^{1/2}$  values.

We also obtain the numerical results from the simulation by calculating the average, the minimum, the maximum, their percentage difference compared with the standard value, and the standard deviation. The results are shown in Table 4. The

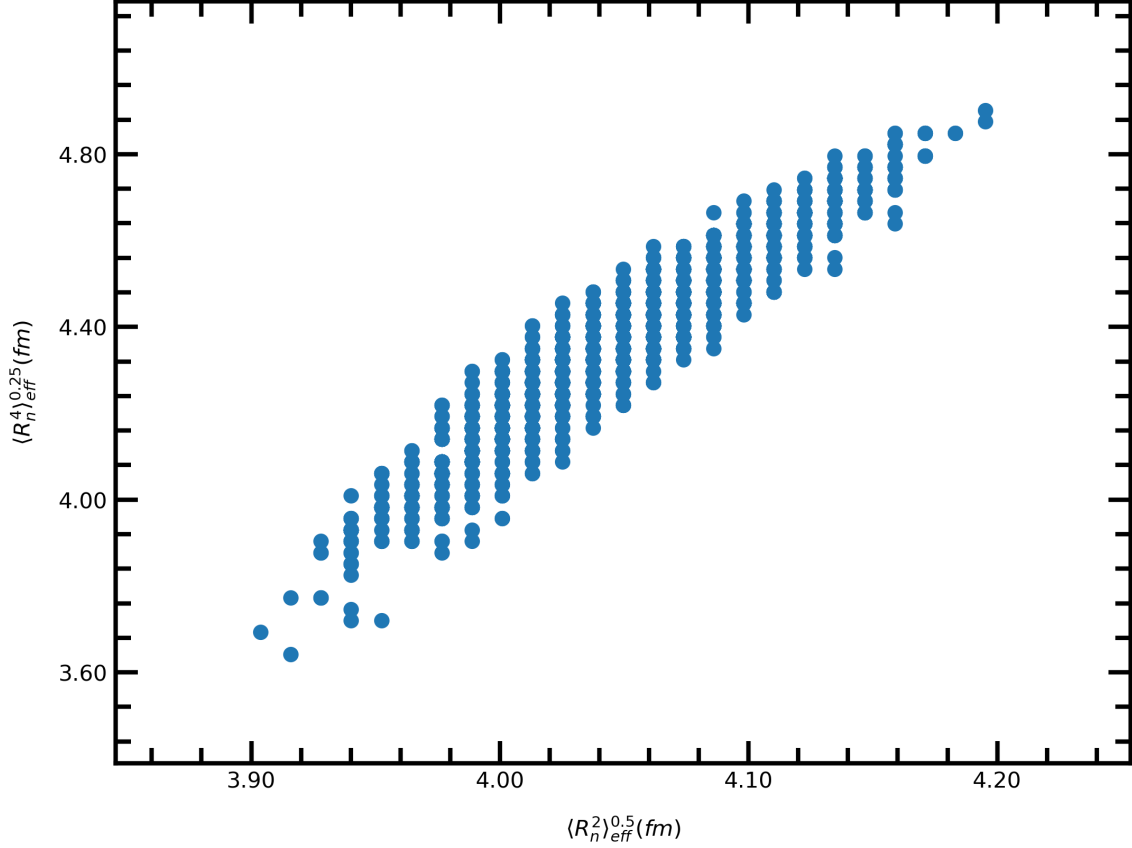


Figure 14: 5000  $\langle R_n^2 \rangle$  and the  $\langle R_n^4 \rangle$  pairs from the Monte-Carlo Simulation. Assume that the neutrino flux is  $3 \times 10^7$  neutrinos per second per cm squared and the detector is filled with 10 tonnes of germanium with 100% detecting efficiency. The scattering events are calculated using the truncated series from Taylor expansion and ignoring the proton and cross terms.

data shows that if the detecting efficiency is constantly 100%, then the measurement of the effective second moment can give us the result that is within  $\pm 0.05\%$  of the standard value while the measurement of the effective fourth moment gives us the result within  $\pm 0.2\%$  of the standard value. Comparing our result with Table 5, the result from Ref. [2] that uses 10 keV bins instead of 5 keV bins, we can see that our  $\langle R_n^4 \rangle^{1/4}$  average value is a lot closer to the standard value, which means that the measurement would give a more accurate result. The comparison shows that the smaller bins can generate higher resolutions, leading to more accurate measurements and results.



$Ge$	SkM* values	Mean	Max	Min	Standard Deviation
$\langle R_n^2 \rangle^{1/2}$ (fm)	4.0495	4.0505	4.1953	3.9037	0.0433
% Difference		0.025	3.57	-3.62	
$\langle R_n^4 \rangle^{1/4}$ (fm)	4.3765	4.3726	4.9017	3.6412	0.1806
% Difference		-0.089	12.1	-16.7	

Table 4: Numerical results for detectors filled with 10 tonnes of natural germanium. Assume that the detecting efficiency is always 100%. The first row is the standard value, minimum, maximum, the percentage differences from the standard value, and the standard deviation for  $\langle R_n^2 \rangle^{1/2}$ . The second row shows the same results for  $\langle R_n^4 \rangle^{1/4}$ .

$Ge$	SkM* values	Mean	Max	Min
$\langle R_n^2 \rangle^{1/2}$ (fm)	4.0495	4.0516	4.2697	3.8792
% Difference		0.05	5	-4
$\langle R_n^4 \rangle^{1/4}$ (fm)	4.3765	4.3603	5.0096	3.7276
% Difference		-0.4	15	-15
Fixing detector efficiency	SkM* values	Mean	Max	Min
$\langle R_n^2 \rangle^{1/2}$ (fm)	4.0495	4.0491	4.1175	3.9857
% Difference		-0.009	1.7	-1.6
$\langle R_n^4 \rangle^{1/4}$ (fm)	4.3765	4.3679	4.6546	4.0826
% Difference		-0.2	7	-7

Table 5: Same as Table. 4. Results from Ref. [2] using the 1.5 tonne germanium detector. The first two rows show the results for allowing the detector efficiency to vary by 10%. The last two rows fixed the detector efficiency.

### 3.4 Future Work

Continuing with our research, there are a few directions that still require more time and effort to investigate. First, in the flux, we only take into account two types of neutrinos, the electron neutrinos and the muon antineutrinos. There is another mono-energetic source that we have not added into our calculation, which gives a constant flux of 29.9 MeV and can make our results more accurate. Second, we have only worked with argon and germanium and only run Monte Carlo simulations with germanium. A good idea would be to add another element like xenon, run Monte Carlo simulations with each, and compare and analyze the trends and differences in their results. Third, during the Monte Carlo simulations, we have been keeping the detecting efficiency at 100%, which is an ideal assumption that is rarely achieved in reality. Varying the detecting efficiency and analyzing its impact on the results for our Monte-Carlo simulations can give us a direct indication of how the consistency of the detectors' performances influences our measurements. Fourth, we used the 5 keV bins which appears to generate better results than the previous implementations using 10 keV bins. It's worth further investigating to run simulations with even lower energy bins like 2 keV bins and compare the results to see if a higher energy resolution can lead to even better measurements. Last, we have calculated the scattering events using the full form factor that adds the proton and cross terms but have not continue our research with this. As shown in Fig. 10, adding the extra terms will slightly reduce the events occurred at lower energies. Therefore, using the full factor and add proton's effective moments as new parameters in the Monte Carlo simulation will give us more accurate results in measuring neutron radius and neutron density distributions.

## 4 Conclusion

We have proven that the neutrino-nucleus coherent elastic scattering can help determine both the second moment and the fourth moment, which also provides information on the neutron radius. By comparing the results of the scattering simulations using the Helm form factor and the first two terms of the form factor expansion for Ar, we proved the applicability of the implementation of the truncated form factor series instead of the full version. Then, by simulating the scattering events using the truncated form factor for a detector filled with 10 tonnes of natural germanium and a neutrino flux of  $3 \times 10^7$  neutrinos per second per  $\text{cm}^2$ , running the Monte Carlo simulation using a bin size of 5 keV, and comparing the results with the previous simulations that used 10 keV bin size, we conclude that if the detector efficiency is fixed, then the CE $\nu$ NS results can determine the neutron radius and the second and fourth moment with less than 1% uncertainty. We also conclude that the smaller energy bins and higher energy resolutions for detectors reduce the uncertainty in the neutron form factor and yield a more accurate measurement of the second and fourth moments. As mentioned previously, future work is needed to make the simulations more realistic and put limits on what is needed and what is possible for measurements.

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## Appendix A Python Code for the Simulations

Starting on the next page, this section will show all the python codes developed in this research that are used to simulate the coherent elastic neutrino-nucleus scattering events and run Monte Carlo simulations. There are a total of 9 files and each file contains detailed description of its purpose.

```

1  ## Hongyong Zhang
2  ## Compute the two types of flux.
3  ## 05/19/2020
4  import math
5  import numpy as np
6  import matplotlib as mpl
7  import matplotlib.pyplot as plt
8  from matplotlib import rc
9  from matplotlib.ticker import (MultipleLocator, FormatStrFormatter, AutoMinorLocator)
10
11 ## Part1: Constants and Calculations
12 #####
13 m = 105.6583745
14 Ev = np.arange(0,m/2,0.5)
15 fe = 96/m**4*(m*Ev**2-2*Ev**3)
16 fu = 16/m**4*(3*m*Ev**2-4*Ev**3)
17
18 ## Part2: Plot
19 #####
20 plt.rcParams['font.size'] = 10
21 plt.rcParams['axes.linewidth'] = 2
22 fig = plt.figure()
23 ax = fig.add_axes([0, 0, 1, 1])
24 # Edit the major and minor ticks of the x and y axes
25 ax.xaxis.set_tick_params(which='major', size=10, width=2, direction='in', top='on')
26 ax.xaxis.set_tick_params(which='minor', size=7, width=2, direction='in', top='on')
27 ax.yaxis.set_tick_params(which='major', size=10, width=2, direction='in', right='on')
28 ax.yaxis.set_tick_params(which='minor', size=7, width=2, direction='in', right='on')
29
30 ax.plot(Ev,fe,label='Electron Neutrinos')
31 ax.plot(Ev,fu,label='Muon Antineutrinos')
32
33 # Set the axis limits
34 ax.set_xlim(0, 55)
35 ax.set_ylim(0, 0.04)
36 # Edit the major and minor tick locations
37 ax.xaxis.set_major_locator(MultipleLocator(10))
38 ax.xaxis.set_major_formatter(FormatStrFormatter('%d'))
39 ax.xaxis.set_minor_locator(AutoMinorLocator(3))
40 ax.yaxis.set_major_locator(MultipleLocator(0.01))
41 ax.yaxis.set_major_formatter(FormatStrFormatter('%.2f'))
42 ax.yaxis.set_minor_locator(AutoMinorLocator(3))
43
44 # Add legend to plot
45 ax.legend(bbox_to_anchor=(0.05, 0.95),loc=2, frameon=False, fontsize=10)
46 ax.set_xlabel('Neutrino Energy (MeV)', labelpad=10)
47 ax.set_ylabel('Probability Density', labelpad=10)
48 plt.savefig('Flux.png', dpi=300, transparent=False, bbox_inches='tight')
49 plt.show()

```

```

1  ## Hongyong Zhang
2  ## Calculate the Helm form factor for different R values.
3  ## 05/19/2020
4  import math
5  import numpy as np
6  import matplotlib as mpl
7  import matplotlib.pyplot as plt
8  from matplotlib import rc
9  from matplotlib.ticker import (MultipleLocator, FormatStrFormatter, AutoMinorLocator)
10
11 ## Part1: Calculations
12 #####
13 s = 0.5
14 def fourier_transform(R0):
15     Q = np.arange(0, 1, 0.01)
16     a = np.sin(Q*R0)/((Q*R0)**2)
17     b = np.cos(Q*R0)/(Q*R0)
18     j1 = a - b
19     e = np.exp(-0.5*((Q*s)**2))
20     F = 3*j1/(Q*R0)*e
21     return (Q,F)
22
23
24 ## Part2: Plot
25 #####
26 plt.rcParams['font.size'] = 10
27 plt.rcParams['axes.linewidth'] = 2
28 fig = plt.figure()
29 ax = fig.add_axes([0, 0, 1, 1])
30 # Edit the major and minor ticks of the x and y axes
31 ax.xaxis.set_tick_params(which='major', size=10, width=2, direction='in', top='on')
32 ax.xaxis.set_tick_params(which='minor', size=7, width=2, direction='in', top='on')
33 ax.yaxis.set_tick_params(which='major', size=10, width=2, direction='in', right='on')
34 ax.yaxis.set_tick_params(which='minor', size=7, width=2, direction='in', right='on')
35 # Set the axis limits
36 ax.set_xlim(0, 1)
37 ax.set_ylim(0, 1)
38 # Edit the major and minor tick locations
39 ax.xaxis.set_major_locator(MultipleLocator(0.2))
40 ax.xaxis.set_major_formatter(FormatStrFormatter('%.1f'))
41 ax.xaxis.set_minor_locator(AutoMinorLocator(3))
42 ax.yaxis.set_major_locator(MultipleLocator(0.2))
43 ax.yaxis.set_major_formatter(FormatStrFormatter('%.1f'))
44 ax.yaxis.set_minor_locator(AutoMinorLocator(3))
45
46
47 (Q,F) = fourier_transform(3)
48 ax.plot(Q,F,label='$R_0 = 3fm$')
49 (Q,F) = fourier_transform(4)
50 ax.plot(Q,F,label='$R_0 = 4fm$')
51 (Q,F) = fourier_transform(5)
52 ax.plot(Q,F,label='$R_0 = 5fm$')
53
54 # Add legend to plot
55 ax.legend(bbox_to_anchor=(0.95, 0.95),loc=1, frameon=False, fontsize=10)
56 ax.set_xlabel('Momentum Transfer Q', labelpad=10)
57 ax.set_ylabel('Form Factor', labelpad=10)
58 plt.savefig('Form_factor.png', dpi=300, transparent=False, bbox_inches='tight')
59 plt.show()

```



```

1  ## Hongyong Zhang
2  ## Calculate the Helm form factor for Ar.
3  ## 05/19/2020
4  import math
5  import numpy as np
6  import matplotlib as mpl
7  import matplotlib.pyplot as plt
8  from matplotlib import rc
9  from matplotlib.ticker import (MultipleLocator, FormatStrFormatter, AutoMinorLocator)
10 import scipy.integrate
11
12 ## Part1: Calculations
13 #####
14 s = 0.5
15 Q = np.arange(0.01, 1, 0.01 )
16 def fourier_transform(R0):
17     a = np.sin(Q*R0)/((Q*R0)**2)
18     b = np.cos(Q*R0)/(Q*R0)
19     j1 = a - b
20     e = np.exp(-0.5*((Q*s)**2))
21     F = 3*j1/(Q*R0)*e
22
23     return F
24
25 # Calculate the full form factor for argon
26 Qw = 22 - 18*(1-4*0.231)
27 ff = fourier_transform(3.43)
28 full_ff = 1/Qw*(22*ff-18*(1-4*0.231)*ff)
29
30
31 ## Part2: Plot
32 #####
33 plt.rcParams['font.size'] = 10
34 plt.rcParams['axes.linewidth'] = 2
35 fig = plt.figure()
36 ax = fig.add_axes([0, 0, 1, 1])
37 # Edit the major and minor ticks of the x and y axes
38 ax.xaxis.set_tick_params(which='major', size=10, width=2, direction='in', top='on')
39 ax.xaxis.set_tick_params(which='minor', size=7, width=2, direction='in', top='on')
40 ax.yaxis.set_tick_params(which='major', size=10, width=2, direction='in', right='on')
41 ax.yaxis.set_tick_params(which='minor', size=7, width=2, direction='in', right='on')
42
43
44 ax.plot(Q, full_ff)
45
46 # Set the axis limits
47 ax.set_xlim(0, 1.1)
48 ax.set_ylim(0, 1.1)
49 # Edit the major and minor tick locations
50 ax.xaxis.set_major_locator(MultipleLocator(0.2))
51 ax.xaxis.set_major_formatter(FormatStrFormatter('%.1f'))
52 ax.xaxis.set_minor_locator(AutoMinorLocator(3))
53 ax.yaxis.set_major_locator(MultipleLocator(0.2))
54 ax.yaxis.set_major_formatter(FormatStrFormatter('%.1f'))
55 ax.yaxis.set_minor_locator(AutoMinorLocator(3))
56
57 ax.set_xlabel('Momentum Transfer Q', labelpad=10)
58 ax.set_ylabel('Form Factor', labelpad=10)
59 plt.savefig('Form_factor_Argon.png', dpi=300, transparent=False, bbox_inches='tight')
60 plt.show()

```

```

1  ## Hongyong Zhang
2  ## Scattering Events for Ar using the Helm form factor.
3  ## 05/19/2020
4  import matplotlib as mpl
5  import matplotlib.pyplot as plt
6  from matplotlib import rc
7  from matplotlib.ticker import (MultipleLocator, FormatStrFormatter, AutoMinorLocator)
8  import numpy as np
9  import pandas as pd
10 import scipy.integrate
11 from numpy import cos, sin, exp, sqrt
12
13 ## Part1: Constants
14 #####
15 ## Mass are in Mev/c^2
16 M = 37215.8
17 m = 105.6583745
18 hbar = 6.582119569e-22
19 c = 3e8
20 event = []
21 energy = []
22
23 ## Part2: Integrations
24 #####
25 for i in range(1,151):
26     t = 0.001*i
27     def integrand1(z):
28         # Calculate dN/dt numerically, variable z is the energy E.
29         s = 0.5e-15
30         R0 = 3.43e-15
31
32
33         Q = (2*(z**2)*t*M/(z**2-z*t))**0.5
34         a = sin(Q*R0/(hbar*c))/((Q*R0/(hbar*c))**2)
35         b = cos(Q*R0/(hbar*c))/(Q*R0/(hbar*c))
36         j1 = a - b
37         e = exp(-0.5*((Q*s/(hbar*c))**2))
38
39         ## form factor
40         F = 3*j1/(Q*R0/(hbar*c))*e
41         Qw = 22 - 18*(1-4*0.231)
42         ## Fermi Constant is in MeV^(-2)
43         Gf = 1.1663787e-5/(1000**2)
44
45         ## full form factor, cross section, and flux
46         full_ff = 1/Qw*(22*F-18*(1-4*0.231)*F)
47         Cross_Section = (Gf**2)/(2*3.1415926)*M*(2-2*t/z+(t/z)**2-
M*t/(z**2))*Qw**2/4*(full_ff**2)
48         fu = 16/m**4*(3*m*z**2-4*z**3)
49         fe = 96/m**4*(m*z**2-2*z**3)
50
51         ## The full integral,multiply by (hbar*c)^2 to normalize.
52         Nt = 54.63e28/39.962383
53         C = 10**11
54         f1 = (hbar*c)**2*C*fu*Cross_Section*(3.154e7)*Nt/1000
55         return f1
56
57     def integrand2(z):
58         # Calculate dN/dt numerically, variable z is the energy E.
59         s = 0.5e-15

```

```

60     R0 = 3.43e-15
61
62
63     Q = (2*(z**2)*t*M/(z**2-z*t))**0.5
64     a = sin(Q*R0/(hbar*c))/((Q*R0/(hbar*c))**2)
65     b = cos(Q*R0/(hbar*c))/(Q*R0/(hbar*c))
66     j1 = a - b
67     e = exp(-0.5*((Q*s/(hbar*c))**2))
68
69     ## form factor
70     F = 3*j1/(Q*R0/(hbar*c))*e
71     Qw = 22 - 18*(1-4*0.231)
72     ## Fermi Constant is in MeV^(-2)
73     Gf = 1.1663787e-5/(1000**2)
74
75     ## full form factor, cross section, and flux
76     full_ff = 1/Qw*(22*F-18*(1-4*0.231)*F)
77     Cross_Section = (Gf**2)/(2*3.1415926)*M*(2-2*t/z+(t/z)**2-
M*t/(z**2))*Qw**2/4*(full_ff**2)
78     fu = 16/m**4*(3*m*z**2-4*z**3)
79     fe = 96/m**4*(m*z**2-2*z**3)
80
81     ## The full integral,multiply by (hbar*c)^2 to normalize.
82     Nt = 54.63e28/39.962383
83     C = 10**11
84     f2 = (hbar*c)**2*C*fe*Cross_Section*(3.154e7)*Nt/1000
85     return f2
86
87
88     I1 = scipy.integrate.quad(integrand1,(t+sqrt(t**2+2*t*M))/2,m/2)
89     I2 = scipy.integrate.quad(integrand2,(t+sqrt(t**2+2*t*M))/2,m/2)
90     event.append(I1[0]+I2[0])
91     energy.append(t*1000)
92
93
94     ## Part3: Plots and Tables
95     #####
96     ## Plot:
97     plt.rcParams['font.size'] = 10
98     plt.rcParams['axes.linewidth'] = 2
99     fig = plt.figure()
100    ax = fig.add_axes([0, 0, 1, 1])
101    # Edit the major and minor ticks of the x and y axes
102    ax.xaxis.set_tick_params(which='major', size=10, width=2, direction='in', top='on')
103    ax.xaxis.set_tick_params(which='minor', size=7, width=2, direction='in', top='on')
104    ax.yaxis.set_tick_params(which='major', size=10, width=2, direction='in', right='on')
105    ax.yaxis.set_tick_params(which='minor', size=7, width=2, direction='in', right='on')
106
107    ax.plot(energy,event)
108
109    # Set the axis limits
110    ax.set_xlim(0, 150)
111    ax.set_ylim(0, 600)
112    # Edit the major and minor tick locations
113    ax.xaxis.set_major_locator(MultipleLocator(15))
114    ax.xaxis.set_major_formatter(FormatStrFormatter('%d'))
115    ax.xaxis.set_minor_locator(AutoMinorLocator(3))
116    ax.yaxis.set_major_locator(MultipleLocator(100))
117    ax.yaxis.set_major_formatter(FormatStrFormatter('%d'))
118    ax.yaxis.set_minor_locator(AutoMinorLocator(3))

```

```
119
120 ax.set_xlabel('Nucleus Recoil Energy (keV)', labelpad=10)
121 ax.set_ylabel('Events/(keV year tonne)', labelpad=10)
122 plt.savefig('Scattering_Events_Ar_Amanik.png', dpi=300, transparent=False,
123             bbox_inches='tight')
124 plt.show()
125
126 table = pd.DataFrame({ 'Bin Range(keV)': energy, 'Events' : event})
127 #print(event1)
128 table.to_csv('table_Ar.csv', index=False, encoding='utf-8')
```

```

1  ## Hongyong Zhang
2  ## Scattering Events for Ar using the first two terms of the form factor expansion.
3  ## 05/19/2020
4  import matplotlib as mpl
5  import matplotlib.pyplot as plt
6  from matplotlib import rc
7  from matplotlib.ticker import (MultipleLocator, FormatStrFormatter, AutoMinorLocator)
8  import numpy as np
9  import scipy.integrate
10 import pandas as pd
11 from numpy import cos, sin, exp, sqrt
12
13 ## Part1: Constants
14 #####
15 ## Mass are in MeV/c^2
16 M = 37215.8
17 m = 105.6583745
18 hbar = 6.582119569e-22
19 c = 3e8
20 R2 = (3.4168e-15)**2
21 R4 = (3.7233e-15)**4
22
23
24 ## Part2: Integrations
25 #####
26 def dNdt(R2,R4):
27     event = []
28     for p in range(1,31):
29         ssum = 0
30         for i in range(1,6):
31             sum = 0
32             t = ((p-1)*5+i)*0.001
33             def integrand1(z,R2,R4):
34                 # Calculate dN/dt numerically, variable z is the energy E.
35                 s = 0.5e-15
36                 R0 = 3.43e-15
37                 Qw = 22 - 18*(1-4*0.231)
38                 Q = (2*(z**2)*t*M/(z**2-z*t))**0.5
39
40
41                 ## full form factor using expansions
42                 full_ff = 22*(1-(Q/(hbar*c))**2/6*R2+(Q/(hbar*c))**4/120*R4)/Qw
43
44                 ## Fermi Constant is in MeV^(-2)
45                 Gf = 1.1663787e-5/(1000**2)
46
47                 ## cross section and flux
48                 Cross_Section = (Gf**2)/(2*3.1415926)*M*(2-2*t/z+(t/z)**2-
M*t/(z**2))*Qw**2/4*(full_ff**2)
49                 fu = 16/m**4*(3*m*z**2-4*z**3)
50                 fe = 96/m**4*(m*z**2-2*z**3)
51
52                 ## The full integral,multiply by (hbar*c)^2 to normalize.
53                 Nt = 54.63e28/39.962383
54                 C = 10**11
55                 f1 = (hbar*c)**2*C*fu*Cross_Section*(3.154e7)*Nt/1000
56                 return f1
57
58             def integrand2(z,R2,R4):
59                 # Calculate dN/dt numerically, variable z is the energy E.

```

```

60     s = 0.5e-15
61     R0 = 3.43e-15
62     Qw = 22 - 18*(1-4*0.231)
63     Q = (2*(z**2)*t*M/(z**2-z*t))**0.5
64
65
66     ## full form factor using expansions
67     full_ff = 22*(1-(Q/(hbar*c))**2/6*R2+(Q/(hbar*c))**4/120*R4)/Qw
68
69     ## Fermi Constant is in MeV^(-2)
70     Gf = 1.1663787e-5/(1000**2)
71
72     ## cross section and flux
73     Cross_Section = (Gf**2)/(2*3.1415926)*M*(2-2*t/z+(t/z)**2-
M*t/(z**2))*Qw**2/4*(full_ff**2)
74     fu = 16/m**4*(3*m*z**2-4*z**3)
75     fe = 96/m**4*(m*z**2-2*z**3)
76
77     ## The full integral,multiply by (hbar*c)^2 to normalize.
78     Nt = 54.63e28/39.962383
79     C = 10**11
80     f2 = (hbar*c)**2*C*fe*Cross_Section*(3.154e7)*Nt/1000
81     return f2
82
83
84     I1 = scipy.integrate.quad(integrand1,(t+sqrt(t**2+2*t*M))/2,m/2,args=
(R2,R4))
85     I2 = scipy.integrate.quad(integrand2,(t+sqrt(t**2+2*t*M))/2,m/2,args=
(R2,R4))
86     sum = I1[0]+I2[0]
87     ssum = ssum + I1[0]+I2[0]
88     #event.append(int(sum))
89     event.append(int(ssum))
90     return event
91
92
93 ## Part3: Plots and Tables
94 #####
95 energy = np.linspace(5,150,num=30)
96 event1 = dNdt(R2,R4)
97 event2 = dNdt(R2*0.9**2,R4*0.9**4)
98 event3 = dNdt(R2*0.8**2,R4*0.8**4)
99 event4 = dNdt(R2*1.1**2,R4*1.1**4)
100 event5 = dNdt(R2*1.2**2,R4*1.2**4)
101
102 table = pd.DataFrame({ 'Bin Range(keV)': energy, 'Events' : event1, 'Events (-10%)' :
event2, 'Events (-20%)' : event3, 'Events (+10%)' : event4, 'Events (+20%)' :
event5})
103 #print(event1)
104 #table.to_csv('table_Ar_Expanded.csv', index=False, encoding='utf-8')
105
106 # # Plot:
107 # plt.rcParams['font.size'] = 10
108 # plt.rcParams['axes.linewidth'] = 2
109 # fig = plt.figure()
110 # ax = fig.add_axes([0, 0, 1, 1])
111 # # Edit the major and minor ticks of the x and y axes
112 # ax.xaxis.set_tick_params(which='major', size=10, width=2, direction='in', top='on')
113 # ax.xaxis.set_tick_params(which='minor', size=7, width=2, direction='in', top='on')

```

```

114 # ax.yaxis.set_tick_params(which='major', size=10, width=2, direction='in',
    right='on')
115 # ax.yaxis.set_tick_params(which='minor', size=7, width=2, direction='in',
    right='on')
116
117 # ax.plot(energy,event1,label='R')
118 # ax.plot(energy,event2,label='0.9R')
119 # ax.plot(energy,event3,label='0.8R')
120 # ax.plot(energy,event4,label='1.1R')
121 # ax.plot(energy,event5,label='1.2R')
122
123 # # Set the axis limits
124 # ax.set_xlim(0, 150)
125 # ax.set_ylim(0, 650)
126 # # Edit the major and minor tick locations
127 # ax.xaxis.set_major_locator(MultipleLocator(20))
128 # ax.xaxis.set_major_formatter(FormatStrFormatter('%d'))
129 # ax.xaxis.set_minor_locator(AutoMinorLocator(3))
130 # ax.yaxis.set_major_locator(MultipleLocator(100))
131 # ax.yaxis.set_major_formatter(FormatStrFormatter('%d'))
132 # ax.yaxis.set_minor_locator(AutoMinorLocator(3))
133
134 # # Add legend to plot
135 # ax.legend(bbox_to_anchor=(0.95, 0.95),loc=1, frameon=False, fontsize=10)
136 # ax.set_xlabel('Nucleus Recoil Energy (keV)', labelpad=10)
137 # ax.set_ylabel('Events/(keV year tonne)', labelpad=10)
138 # plt.savefig('Scattering_Events_Ar.png', dpi=300, transparent=False,
    bbox_inches='tight')
139 # plt.show()
140
141 # ## Second Plot: Compare the difference between this and the Amanik's form factor.
142 # (energy0,event0) = np.loadtxt('table_Ar.csv', unpack=True, delimiter=',',
    skiprows=1)
143 # diff = event1-event0
144 # plt.rcParams['font.size'] = 10
145 # plt.rcParams['axes.linewidth'] = 2
146 # fig = plt.figure()
147 # ax2 = fig.add_axes([0, 0, 1, 0.4])
148 # ax1 = fig.add_axes([0, 0.6, 1, 0.4])
149
150 # # # Set the axis limits
151 # ax1.set_xlim(0, 150)
152 # ax1.set_ylim(0, 650)
153 # ax2.set_xlim(0, 150)
154 # ax2.set_ylim(-10, 80)
155 # # Edit the major and minor ticks of the x and y axes
156 # ax1.xaxis.set_tick_params(which='major', size=10, width=2, direction='in',
    top='on')
157 # ax1.xaxis.set_tick_params(which='minor', size=7, width=2, direction='in', top='on')
158 # ax1.yaxis.set_tick_params(which='major', size=10, width=2, direction='in',
    right='on')
159 # ax1.yaxis.set_tick_params(which='minor', size=7, width=2, direction='in',
    right='on')
160 # ax2.xaxis.set_tick_params(which='major', size=10, width=2, direction='in',
    top='on')
161 # ax2.xaxis.set_tick_params(which='minor', size=7, width=2, direction='in', top='on')
162 # ax2.yaxis.set_tick_params(which='major', size=10, width=2, direction='in',
    right='on')
163 # ax2.yaxis.set_tick_params(which='minor',43 size=7, width=2, direction='in',
    right='on')

```

```

164
165 # ax1.plot(energy,event1,label='With Expansion')
166 # ax1.plot(energy,event0,label='Without Expansion')
167 # ax2.plot(energy,diff)
168
169 # # Edit the major and minor tick locations
170 # ax1.xaxis.set_major_locator(MultipleLocator(20))
171 # ax1.xaxis.set_major_formatter(FormatStrFormatter('%d'))
172 # ax1.xaxis.set_minor_locator(AutoMinorLocator(3))
173 # ax1.yaxis.set_major_locator(MultipleLocator(100))
174 # ax1.yaxis.set_major_formatter(FormatStrFormatter('%d'))
175 # ax1.yaxis.set_minor_locator(AutoMinorLocator(3))
176 # ax2.xaxis.set_major_locator(MultipleLocator(20))
177 # ax2.xaxis.set_major_formatter(FormatStrFormatter('%d'))
178 # ax2.xaxis.set_minor_locator(AutoMinorLocator(3))
179 # ax2.yaxis.set_major_locator(MultipleLocator(10))
180 # ax2.yaxis.set_major_formatter(FormatStrFormatter('%d'))
181 # ax2.yaxis.set_minor_locator(AutoMinorLocator(3))
182 # # Add legend to plot
183 # ax1.legend(bbox_to_anchor=(0.95, 0.95),loc=1, frameon=False, fontsize=10)
184 # ax2.set_xlabel('Nucleus Recoil Energy (keV)', labelpad=10)
185 # ax2.set_ylabel('Difference in Events/(keV year tonne)', labelpad=10)
186 # ax1.set_ylabel('Events/(keV year tonne)', labelpad=10)
187 # plt.savefig('Scattering_Events_Ar_diff.png', dpi=300, transparent=False,
188 #             bbox_inches='tight')
189 # plt.show()
190
191 ## Bar Plot:
192 plt.rcParams['font.size'] = 10
193 plt.rcParams['axes.linewidth'] = 2
194 fig = plt.figure()
195 ax = fig.add_axes([0, 0, 1, 1])
196 # Edit the major and minor ticks of the x and y axes
197 ax.xaxis.set_tick_params(which='major', size=2, width=2, direction='in', top='on')
198 ax.yaxis.set_tick_params(which='major', size=5, width=2, direction='in', right='on')
199 ax.yaxis.set_tick_params(which='minor', size=2, width=2, direction='in', right='on')
200
201 # Set the axis limits
202 ax.set_xlim(0, 150)
203 ax.set_ylim(0, 3200)
204 # Edit the major and minor tick locations
205 ax.yaxis.set_major_locator(MultipleLocator(800))
206 ax.yaxis.set_major_formatter(FormatStrFormatter('%d'))
207 ax.yaxis.set_minor_locator(AutoMinorLocator(2))
208
209 plt.bar(energy,event1,align='center',width=4)
210 plt.xticks(energy)
211 plt.xticks(rotation=90)
212 plt.ylabel("Number of Counts (keV year tonne)")
213 plt.xlabel("Nucleus Recoil Energy (keV)")
214 # for i, v in enumerate(event1):
215 #     plt.text(energy[i] -0.25, v , str(v))
216 plt.savefig('Scattering_Events_Ar_binned.png', dpi=300, transparent=False,
217 #             bbox_inches='tight')
218 plt.show()
219

```



```

1  ## Hongyong Zhang
2  ## Scattering Events for Ge using the first two terms of the form factor expansion
   and
3  ## only the neutron form factor.
4  ## 05/19/2020
5  import matplotlib as mpl
6  import matplotlib.pyplot as plt
7  from matplotlib import rc
8  from matplotlib.ticker import (MultipleLocator, FormatStrFormatter, AutoMinorLocator)
9  import numpy as np
10 import scipy.integrate
11 import pandas as pd
12 from numpy import cos, sin, exp, sqrt
13
14 # PART 1: CONSTANTS
15 #####
16 ## Mass are in MeV/c^2
17 M_bar =
   (69.9242474*0.2057+71.92207589*0.2745+72.9234589*0.0775+73.9211778*0.3650+75.9214026*
   0.0773)*931.5
18 m = 105.6583745
19 hbar = 6.582119569e-22
20 c = 3e8
21 R2 = (4.0495e-15)**2
22 R4 = (4.3765e-15)**4
23 M_detector = 1000*5.5866e29
24 Na = 6.022e23
25
26
27 #PART2: Integration
28 #####
29 def dNdt(R2,R4):
30     event = []
31     event_bin=[]
32     for p in range(1,17):
33
34         sum_bin = 0
35         for i in range(1,6):
36             ssum = 0
37             t = ((p-1)*5+i)*0.001
38             def integrand1(z,R2,R4):
39                 # Calculate dN/dt numerically, variable z is the energy E.
40                 s = 0.5e-15
41                 Qw = 22 - 18*(1-4*0.231)
42                 ## Fermi Constant is in MeV^(-2)
43                 Gf = 1.1663787e-5/(1000**2)
44                 # Calculate the cross section for Ge. Summing over different
   isotopes.
45                 M1 = 69.9242474*931.5
46                 M2 = 71.92207589*931.5
47                 M3 = 72.9234589*931.5
48                 M4 = 73.9211778*931.5
49                 M5 =75.9214026*931.5
50                 X1 = 0.2057
51                 X2 = 0.2745
52                 X3 = 0.0775
53                 X4 = 0.3650
54                 X5 = 0.0773
55                 N1 = 38
56                 N2 = 40

```

```

57     N3 = 41
58     N4 = 42
59     N5 = 44
60     Q = (2*(z**2)*t*M_bar/(z**2-z*t))**0.5/(hbar*c)
61     R2_eff = R2
62     R4_eff = R4
63
64     cs1 = Gf**2/(8*3.1415926)*(X1*N1**2*(2-2*t/z+(t/z)**2)-X1*N1**2*M1*
(t/z**2)-R2_eff*X1*N1**2*M1*(2-2*t/z+(t/z)**2)*Q**2/(3*M_bar)+R2_eff*X1*N1**2*M1**2*
(t/z**2)*Q**2/(3*M_bar)+R2_eff**2*X1*N1**2*M1**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff**2*X1*N1**2*M1**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff*X1*N1**2*M1**2*(2-2*t/z+
(t/z)**2)*Q**4/(60.0*M_bar**2)-R4_eff*X1*N1**2*M1**3*(t/z**2)*Q**4/(60.0*M_bar**2)-
R2_eff*R4_eff*X1*N1**2*M1**3*(2-2*t/z+
(t/z)**2)*Q**6/(360.0*M_bar**3)+R2_eff*R4_eff*X1*N1**2*M1**4*
(t/z**2)*Q**6/(360.0*M_bar**3))
65     cs2 = Gf**2/(8*3.1415926)*(X2*N2**2*(2-2*t/z+(t/z)**2)-X2*N2**2*M2*
(t/z**2)-R2_eff*X2*N2**2*M2*(2-2*t/z+(t/z)**2)*Q**2/(3*M_bar)+R2_eff*X2*N2**2*M2**2*
(t/z**2)*Q**2/(3*M_bar)+R2_eff**2*X2*N2**2*M2**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff**2*X2*N2**2*M2**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff*X2*N2**2*M2**2*(2-2*t/z+
(t/z)**2)*Q**4/(60.0*M_bar**2)-R4_eff*X2*N2**2*M2**3*(t/z**2)*Q**4/(60.0*M_bar**2)-
R2_eff*R4_eff*X2*N2**2*M2**3*(2-2*t/z+
(t/z)**2)*Q**6/(360.0*M_bar**3)+R2_eff*R4_eff*X2*N2**2*M2**4*
(t/z**2)*Q**6/(360.0*M_bar**3))
66     cs3 = Gf**2/(8*3.1415926)*(X3*N3**2*(2-2*t/z+(t/z)**2)-X3*N3**2*M3*
(t/z**2)-R2_eff*X3*N3**2*M3*(2-2*t/z+(t/z)**2)*Q**2/(3*M_bar)+R2_eff*X3*N3**2*M3**2*
(t/z**2)*Q**2/(3*M_bar)+R2_eff**2*X3*N3**2*M3**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff**2*X3*N3**2*M3**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff*X3*N3**2*M3**2*(2-2*t/z+
(t/z)**2)*Q**4/(60.0*M_bar**2)-R4_eff*X3*N3**2*M3**3*(t/z**2)*Q**4/(60.0*M_bar**2)-
R2_eff*R4_eff*X3*N3**2*M3**3*(2-2*t/z+
(t/z)**2)*Q**6/(360.0*M_bar**3)+R2_eff*R4_eff*X3*N3**2*M3**4*
(t/z**2)*Q**6/(360.0*M_bar**3))
67     cs4 = Gf**2/(8*3.1415926)*(X4*N4**2*(2-2*t/z+(t/z)**2)-X4*N4**2*M4*
(t/z**2)-R2_eff*X4*N4**2*M4*(2-2*t/z+(t/z)**2)*Q**2/(3*M_bar)+R2_eff*X4*N4**2*M4**2*
(t/z**2)*Q**2/(3*M_bar)+R2_eff**2*X4*N4**2*M4**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff**2*X4*N4**2*M4**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff*X4*N4**2*M4**2*(2-2*t/z+
(t/z)**2)*Q**4/(60.0*M_bar**2)-R4_eff*X4*N4**2*M4**3*(t/z**2)*Q**4/(60.0*M_bar**2)-
R2_eff*R4_eff*X4*N4**2*M4**3*(2-2*t/z+
(t/z)**2)*Q**6/(360.0*M_bar**3)+R2_eff*R4_eff*X4*N4**2*M4**4*
(t/z**2)*Q**6/(360.0*M_bar**3))
68     cs5 = Gf**2/(8*3.1415926)*(X5*N5**2*(2-2*t/z+(t/z)**2)-X5*N5**2*M5*
(t/z**2)-R2_eff*X5*N5**2*M5*(2-2*t/z+(t/z)**2)*Q**2/(3*M_bar)+R2_eff*X5*N5**2*M5**2*
(t/z**2)*Q**2/(3*M_bar)+R2_eff**2*X5*N5**2*M5**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff**2*X5*N5**2*M5**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff*X5*N5**2*M5**2*(2-2*t/z+
(t/z)**2)*Q**4/(60.0*M_bar**2)-R4_eff*X5*N5**2*M5**3*(t/z**2)*Q**4/(60.0*M_bar**2)-
R2_eff*R4_eff*X5*N5**2*M5**3*(2-2*t/z+
(t/z)**2)*Q**6/(360.0*M_bar**3)+R2_eff*R4_eff*X5*N5**2*M5**4*
(t/z**2)*Q**6/(360.0*M_bar**3))
69     cross_section = cs1 + cs2 + cs3 + cs4 + cs5
70
71     ## flux
72     fu = 16/m**4*(3*m*z**2-4*z**3)
73     fe = 96/m**4*(m*z**2-2*z**3)
74
75     ## The full integral,multiply by (hbar*c)^2 to normalize.
76     C = 10**11

```

```

77         f1 = (hbar*c)**2*C*fu*cross_section*(3.156e7)*M_detector/1000
78         return f1
79
80     def integrand2(z,R2,R4):
81         # Calculate dN/dt numerically, variable z is the energy E.
82         s = 0.5e-15
83         Qw = 22 - 18*(1-4*0.231)
84         ## Fermi Constant is in MeV^(-2)
85         Gf = 1.1663787e-5/(1000**2)
86         # Calculate the cross section for Ge. Summing over different
isotoopes.
87         M1 = 69.9242474*931.5
88         M2 = 71.92207589*931.5
89         M3 = 72.9234589*931.5
90         M4 = 73.9211778*931.5
91         M5 = 75.9214026*931.5
92         X1 = 0.2057
93         X2 = 0.2745
94         X3 = 0.0775
95         X4 = 0.3650
96         X5 = 0.0773
97         N1 = 38
98         N2 = 40
99         N3 = 41
100        N4 = 42
101        N5 = 44
102        M_bar =
(69.9242474*0.2057+71.92207589*0.2745+72.9234589*0.0775+73.9211778*0.3650+75.9214026*
0.0773)*931.5
103        Q = (2*(z**2)*t*M_bar/(z**2-z*t))**0.5/(hbar*c)
104        R2_eff = R2
105        R4_eff = R4
106
107        cs1 = Gf**2/(8*3.1415926)*(X1*N1**2*(2-2*t/z+(t/z)**2)-X1*N1**2*M1*
(t/z**2)-R2_eff*X1*N1**2*M1*(2-2*t/z+(t/z)**2)*Q**2/(3*M_bar)+R2_eff*X1*N1**2*M1**2*
(t/z**2)*Q**2/(3*M_bar)+R2_eff**2*X1*N1**2*M1**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff**2*X1*N1**2*M1**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff*X1*N1**2*M1**2*(2-2*t/z+
(t/z)**2)*Q**4/(60.0*M_bar**2)-R4_eff*X1*N1**2*M1**3*(t/z**2)*Q**4/(60.0*M_bar**2)-
R2_eff*R4_eff*X1*N1**2*M1**3*(2-2*t/z+
(t/z)**2)*Q**6/(360.0*M_bar**3)+R2_eff*R4_eff*X1*N1**2*M1**4*
(t/z**2)*Q**6/(360.0*M_bar**3))
108        cs2 = Gf**2/(8*3.1415926)*(X2*N2**2*(2-2*t/z+(t/z)**2)-X2*N2**2*M2*
(t/z**2)-R2_eff*X2*N2**2*M2*(2-2*t/z+(t/z)**2)*Q**2/(3*M_bar)+R2_eff*X2*N2**2*M2**2*
(t/z**2)*Q**2/(3*M_bar)+R2_eff**2*X2*N2**2*M2**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff**2*X2*N2**2*M2**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff*X2*N2**2*M2**2*(2-2*t/z+
(t/z)**2)*Q**4/(60.0*M_bar**2)-R4_eff*X2*N2**2*M2**3*(t/z**2)*Q**4/(60.0*M_bar**2)-
R2_eff*R4_eff*X2*N2**2*M2**3*(2-2*t/z+
(t/z)**2)*Q**6/(360.0*M_bar**3)+R2_eff*R4_eff*X2*N2**2*M2**4*
(t/z**2)*Q**6/(360.0*M_bar**3))
109        cs3 = Gf**2/(8*3.1415926)*(X3*N3**2*(2-2*t/z+(t/z)**2)-X3*N3**2*M3*
(t/z**2)-R2_eff*X3*N3**2*M3*(2-2*t/z+(t/z)**2)*Q**2/(3*M_bar)+R2_eff*X3*N3**2*M3**2*
(t/z**2)*Q**2/(3*M_bar)+R2_eff**2*X3*N3**2*M3**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff**2*X3*N3**2*M3**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff*X3*N3**2*M3**2*(2-2*t/z+
(t/z)**2)*Q**4/(60.0*M_bar**2)-R4_eff*X3*N3**2*M3**3*(t/z**2)*Q**4/(60.0*M_bar**2)-
R2_eff*R4_eff*X3*N3**2*M3**3*(2-2*t/z+
(t/z)**2)*Q**6/(360.0*M_bar**3)+R2_eff*R4_eff*X3*N3**2*M3**4*
(t/z**2)*Q**6/(360.0*M_bar**3))

```

```

110     cs4 = Gf**2/(8*3.1415926)*(X4*N4**2*(2-2*t/z+(t/z)**2)-X4*N4**2*M4*
(t/z**2)-R2_eff*X4*N4**2*M4*(2-2*t/z+(t/z)**2)*Q**2/(3*M_bar)+R2_eff*X4*N4**2*M4**2*
(t/z**2)*Q**2/(3*M_bar)+R2_eff**2*X4*N4**2*M4**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff**2*X4*N4**2*M4**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff*X4*N4**2*M4**2*(2-2*t/z+
(t/z)**2)*Q**4/(60.0*M_bar**2)-R4_eff*X4*N4**2*M4**3*(t/z**2)*Q**4/(60.0*M_bar**2)-
R2_eff*R4_eff*X4*N4**2*M4**3*(2-2*t/z+
(t/z)**2)*Q**6/(360.0*M_bar**3)+R2_eff*R4_eff*X4*N4**2*M4**4*
(t/z**2)*Q**6/(360.0*M_bar**3))
111     cs5 = Gf**2/(8*3.1415926)*(X5*N5**2*(2-2*t/z+(t/z)**2)-X5*N5**2*M5*
(t/z**2)-R2_eff*X5*N5**2*M5*(2-2*t/z+(t/z)**2)*Q**2/(3*M_bar)+R2_eff*X5*N5**2*M5**2*
(t/z**2)*Q**2/(3*M_bar)+R2_eff**2*X5*N5**2*M5**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff**2*X5*N5**2*M5**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff*X5*N5**2*M5**2*(2-2*t/z+
(t/z)**2)*Q**4/(60.0*M_bar**2)-R4_eff*X5*N5**2*M5**3*(t/z**2)*Q**4/(60.0*M_bar**2)-
R2_eff*R4_eff*X5*N5**2*M5**3*(2-2*t/z+
(t/z)**2)*Q**6/(360.0*M_bar**3)+R2_eff*R4_eff*X5*N5**2*M5**4*
(t/z**2)*Q**6/(360.0*M_bar**3))
112     cross_section = cs1 + cs2 + cs3 + cs4 + cs5
113
114     ## flux
115     fu = 16/m**4*(3*m*z**2-4*z**3)
116     fe = 96/m**4*(m*z**2-2*z**3)
117
118     ## The full integral,multiply by (hbar*c)^2 to normalize.
119     C = 10**11
120     f2 = (hbar*c)**2*C*fe*cross_section*(3.156e7)*M_detector/1000
121     return f2
122
123
124     I1 = scipy.integrate.quad(integrand1,(t+sqrt(t**2+2*t*M_bar))/2,m/2,args=
(R2,R4))
125     I2 = scipy.integrate.quad(integrand2,(t+sqrt(t**2+2*t*M_bar))/2,m/2,args=
(R2,R4))
126     ssum = ssum + I1[0]+I2[0]
127     #print(ssum)
128     sum_bin = sum_bin + ssum
129     event.append(int(ssum))
130
131     event_bin.append(int(sum_bin))
132     return (event,event_bin)
133
134
135     ## Part3: Plots and Tables
136     #####
137     energy = np.linspace(1,80,num=80)
138     energy2 = np.linspace(5,80,num=16)
139     (event1,event_bin1) = dNdt(R2,R4)
140     # (event2,event_bin2) = dNdt(R2*0.9**2,R4*0.9**4)
141     # (event3,event_bin3) = dNdt(R2*0.8**2,R4*0.8**4)
142     # (event4,event_bin4) = dNdt(R2*1.1**2,R4*1.1**4)
143     # (event5,event_bin5) = dNdt(R2*1.2**2,R4*1.2**4)
144     #print("event=",event1)
145     #print("binned event=",event_bin1)
146     table = pd.DataFrame({ 'Bin Range(keV)': energy, 'Events' : event1})
147     #table = pd.DataFrame({ 'Bin Range(keV)': energy, 'Events' : event1, 'Events (-10%)'
: event2, 'Events (-20%)' : event3, 'Events (+10%)' : event4, 'Events (+20%)' :
event5})
148     table.to_csv('table_Ge_data.csv', index=False, encoding='utf-8')
149

```

```

150
151     # plt.plot(energy,event1)
152     # plt.bar(energy2,event_bin1,align='center',width=4)
153     # plt.xticks(energy2)
154     # plt.ylabel("Number of Counts")
155     # plt.xlabel("Energy/keV")
156     # plt.title("# of Events per (keV year tonne) for Ge")
157     # for i, v in enumerate(event_bin1):
158     #     plt.text(energy2[i] -0.25, v , str(v))
159     # plt.show()
160
161
162 ## Bar Plot:
163 plt.rcParams['font.size'] = 10
164 plt.rcParams['axes.linewidth'] = 2
165 fig = plt.figure()
166 ax = fig.add_axes([0, 0, 1, 1])
167 # Edit the major and minor ticks of the x and y axes
168 ax.xaxis.set_tick_params(which='major', size=2, width=2, direction='in', top='on')
169 ax.yaxis.set_tick_params(which='major', size=5, width=2, direction='in', right='on')
170 ax.yaxis.set_tick_params(which='minor', size=2, width=2, direction='in', right='on')
171
172 # Set the axis limits
173 ax.set_xlim(0, 80)
174 ax.set_ylim(0, 11000)
175 # Edit the major and minor tick locations
176 ax.yaxis.set_major_locator(MultipleLocator(1000))
177 ax.yaxis.set_major_formatter(FormatStrFormatter('%d'))
178 ax.yaxis.set_minor_locator(AutoMinorLocator(2))
179
180 plt.bar(energy2,event_bin1,align='center',width=4)
181 plt.xticks(energy2)
182 plt.xticks(rotation=90)
183 plt.ylabel("Number of Counts (keV year tonne)")
184 plt.xlabel("Nucleus Recoil Energy (keV)")
185 for i, v in enumerate(event_bin1):
186     plt.text(energy2[i] -0.25, v , str(v))
187 plt.savefig('Scattering_Events_Ge_binned.png', dpi=300, transparent=False,
188           bbox_inches='tight')
189 plt.show()

```

```

1  ## Hongyong Zhang
2  ## Scattering Events for Ge including the proton and cross terms.
3  ## 05/19/2020
4  import matplotlib as mpl
5  import matplotlib.pyplot as plt
6  from matplotlib import rc
7  from matplotlib.ticker import (MultipleLocator, FormatStrFormatter, AutoMinorLocator)
8  import numpy as np
9  import scipy.integrate
10 import pandas as pd
11 from numpy import cos, sin, exp, sqrt
12
13 ##### Calculate the Contribution from the proton part.
14
15 # PART 1: CONSTANTS
16 #####
17 # Some constants that we need.
18 ## Mass are in MeV/c^2
19 M_bar =
    (69.9242474*0.2057+71.92207589*0.2745+72.9234589*0.0775+73.9211778*0.3650+75.9214026*
    0.0773)*931.5
20 m = 105.6583745
21 hbar = 6.582119569e-22
22 c = 3e8
23 R2 = (4.0495e-15)**2
24 R4 = (4.3765e-15)**4
25 M_detector = 1000*5.5866e29
26 Na = 6.022e23
27
28 s = 0.5e-15
29 #Qw = 22 - 18*(1-4*0.231)
30 ## Fermi Constant is in MeV^(-2)
31 Gf = 1.1663787e-5/(1000**2)
32 # Calculate the cross section for Ge. Summing over different isotopes.
33 M1 = 69.9242474*931.5
34 M2 = 71.92207589*931.5
35 M3 = 72.9234589*931.5
36 M4 = 73.9211778*931.5
37 M5 = 75.9214026*931.5
38 X1 = 0.2057
39 X2 = 0.2745
40 X3 = 0.0775
41 X4 = 0.3650
42 X5 = 0.0773
43 N1 = 38
44 N2 = 40
45 N3 = 41
46 N4 = 42
47 N5 = 44
48 Z = 32
49 s2thw = 2.31e-1
50 alpha = 1-4*s2thw
51
52 R2_eff = R2
53 R4_eff = R4
54
55 ## Calculate the R2_eff and R4_eff for proton
56 R2_eff_p = (X1*M1*(15.4282334386e-30)+X2*M2*(15.8950606333e-30)+X3*M3*
    (15.6649486626e-30)+X4*M4*(15.9736268768e-30)+X5*M5*(16.0256582784e-
    30))/(X1*M1+X2*M2+X3*M3+X4*M4+X5*M5)

```

```

57 R4_eff_p = (X1*M1*(323.6887827706e-60)+X2*M2*(344.1084360256e-60)+X3*M3*
(330.9393314606e-60)+X4*M4*(345.718316624e-60)+X5*M5*(345.8995154439e-
60))/(X1*M1+X2*M2+X3*M3+X4*M4+X5*M5)
58
59
60 #PART2: Integration
61 #####
62 def dNdt(R2,R4):
63     event = []
64     event_bin = []
65     for p in range(1,17):
66
67         sum_bin = 0
68         for i in range(1,6):
69             ssum = 0
70             t = ((p-1)*5+i)*0.001
71             ## The neutron part for flux fu.
72             def integrand1(z,R2,R4):
73                 R2_eff = R2
74                 R4_eff = R4
75                 # Calculate dN/dt numerically, variable z is the energy E.
76                 Q = (2*(z**2)*t*M_bar/(z**2-z*t))**0.5/(hbar*c)
77                 # Five parts of cross section.
78                 cs1 = Gf**2/(8*3.1415926)*(X1*N1**2*(2-2*t/z+(t/z)**2)-X1*N1**2*M1*
(t/z**2)-R2_eff*X1*N1**2*M1*(2-2*t/z+(t/z)**2)*Q**2/(3*M_bar)+R2_eff*X1*N1**2*M1**2*
(t/z**2)*Q**2/(3*M_bar)+R2_eff**2*X1*N1**2*M1**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff**2*X1*N1**2*M1**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff*X1*N1**2*M1**2*(2-2*t/z+
(t/z)**2)*Q**4/(60.0*M_bar**2)-R4_eff*X1*N1**2*M1**3*(t/z**2)*Q**4/(60.0*M_bar**2)-
R2_eff*R4_eff*X1*N1**2*M1**3*(2-2*t/z+
(t/z)**2)*Q**6/(360.0*M_bar**3)+R2_eff*R4_eff*X1*N1**2*M1**4*
(t/z**2)*Q**6/(360.0*M_bar**3))
79                 cs2 = Gf**2/(8*3.1415926)*(X2*N2**2*(2-2*t/z+(t/z)**2)-X2*N2**2*M2*
(t/z**2)-R2_eff*X2*N2**2*M2*(2-2*t/z+(t/z)**2)*Q**2/(3*M_bar)+R2_eff*X2*N2**2*M2**2*
(t/z**2)*Q**2/(3*M_bar)+R2_eff**2*X2*N2**2*M2**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff**2*X2*N2**2*M2**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff*X2*N2**2*M2**2*(2-2*t/z+
(t/z)**2)*Q**4/(60.0*M_bar**2)-R4_eff*X2*N2**2*M2**3*(t/z**2)*Q**4/(60.0*M_bar**2)-
R2_eff*R4_eff*X2*N2**2*M2**3*(2-2*t/z+
(t/z)**2)*Q**6/(360.0*M_bar**3)+R2_eff*R4_eff*X2*N2**2*M2**4*
(t/z**2)*Q**6/(360.0*M_bar**3))
80                 cs3 = Gf**2/(8*3.1415926)*(X3*N3**2*(2-2*t/z+(t/z)**2)-X3*N3**2*M3*
(t/z**2)-R2_eff*X3*N3**2*M3*(2-2*t/z+(t/z)**2)*Q**2/(3*M_bar)+R2_eff*X3*N3**2*M3**2*
(t/z**2)*Q**2/(3*M_bar)+R2_eff**2*X3*N3**2*M3**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff**2*X3*N3**2*M3**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff*X3*N3**2*M3**2*(2-2*t/z+
(t/z)**2)*Q**4/(60.0*M_bar**2)-R4_eff*X3*N3**2*M3**3*(t/z**2)*Q**4/(60.0*M_bar**2)-
R2_eff*R4_eff*X3*N3**2*M3**3*(2-2*t/z+
(t/z)**2)*Q**6/(360.0*M_bar**3)+R2_eff*R4_eff*X3*N3**2*M3**4*
(t/z**2)*Q**6/(360.0*M_bar**3))
81                 cs4 = Gf**2/(8*3.1415926)*(X4*N4**2*(2-2*t/z+(t/z)**2)-X4*N4**2*M4*
(t/z**2)-R2_eff*X4*N4**2*M4*(2-2*t/z+(t/z)**2)*Q**2/(3*M_bar)+R2_eff*X4*N4**2*M4**2*
(t/z**2)*Q**2/(3*M_bar)+R2_eff**2*X4*N4**2*M4**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff**2*X4*N4**2*M4**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff*X4*N4**2*M4**2*(2-2*t/z+
(t/z)**2)*Q**4/(60.0*M_bar**2)-R4_eff*X4*N4**2*M4**3*(t/z**2)*Q**4/(60.0*M_bar**2)-
R2_eff*R4_eff*X4*N4**2*M4**3*(2-2*t/z+
(t/z)**2)*Q**6/(360.0*M_bar**3)+R2_eff*R4_eff*X4*N4**2*M4**4*
(t/z**2)*Q**6/(360.0*M_bar**3))

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82         cs5 = Gf**2/(8*3.1415926)*(X5*N5**2*(2-2*t/z+(t/z)**2)-X5*N5**2*M5*
(t/z**2)-R2_eff*X5*N5**2*M5*(2-2*t/z+(t/z)**2)*Q**2/(3*M_bar)+R2_eff*X5*N5**2*M5**2*
(t/z**2)*Q**2/(3*M_bar)+R2_eff**2*X5*N5**2*M5**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff**2*X5*N5**2*M5**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff*X5*N5**2*M5**2*(2-2*t/z+
(t/z)**2)*Q**4/(60.0*M_bar**2)-R4_eff*X5*N5**2*M5**3*(t/z**2)*Q**4/(60.0*M_bar**2)-
R2_eff*R4_eff*X5*N5**2*M5**3*(2-2*t/z+
(t/z)**2)*Q**6/(360.0*M_bar**3)+R2_eff*R4_eff*X5*N5**2*M5**4*
(t/z**2)*Q**6/(360.0*M_bar**3))
83         cross_section = cs1 + cs2 + cs3 + cs4 + cs5
84
85         ## flux
86         fu = 16/m**4*(3*m*z**2-4*z**3)
87         fe = 96/m**4*(m*z**2-2*z**3)
88
89         ## The full integral,multiply by (hbar*c)^2 to normalize.
90         C = 10**11
91         f1 = (hbar*c)**2*C*fu*cross_section*(3.156e7)*M_detector/1000
92         return f1
93
94         ## The neutron part for flux fe.
95         def integrand2(z,R2,R4):
96             R2_eff = R2
97             R4_eff = R4
98             # Calculate dN/dt numerically, variable z is the energy E.
99             Q = (2*(z**2)*t*M_bar/(z**2-z*t))**0.5/(hbar*c)
100             #Five parts of cross section.
101             cs1 = Gf**2/(8*3.1415926)*(X1*N1**2*(2-2*t/z+(t/z)**2)-X1*N1**2*M1*
(t/z**2)-R2_eff*X1*N1**2*M1*(2-2*t/z+(t/z)**2)*Q**2/(3*M_bar)+R2_eff*X1*N1**2*M1**2*
(t/z**2)*Q**2/(3*M_bar)+R2_eff**2*X1*N1**2*M1**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff**2*X1*N1**2*M1**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff*X1*N1**2*M1**2*(2-2*t/z+
(t/z)**2)*Q**4/(60.0*M_bar**2)-R4_eff*X1*N1**2*M1**3*(t/z**2)*Q**4/(60.0*M_bar**2)-
R2_eff*R4_eff*X1*N1**2*M1**3*(2-2*t/z+
(t/z)**2)*Q**6/(360.0*M_bar**3)+R2_eff*R4_eff*X1*N1**2*M1**4*
(t/z**2)*Q**6/(360.0*M_bar**3))
102             cs2 = Gf**2/(8*3.1415926)*(X2*N2**2*(2-2*t/z+(t/z)**2)-X2*N2**2*M2*
(t/z**2)-R2_eff*X2*N2**2*M2*(2-2*t/z+(t/z)**2)*Q**2/(3*M_bar)+R2_eff*X2*N2**2*M2**2*
(t/z**2)*Q**2/(3*M_bar)+R2_eff**2*X2*N2**2*M2**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff**2*X2*N2**2*M2**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff*X2*N2**2*M2**2*(2-2*t/z+
(t/z)**2)*Q**4/(60.0*M_bar**2)-R4_eff*X2*N2**2*M2**3*(t/z**2)*Q**4/(60.0*M_bar**2)-
R2_eff*R4_eff*X2*N2**2*M2**3*(2-2*t/z+
(t/z)**2)*Q**6/(360.0*M_bar**3)+R2_eff*R4_eff*X2*N2**2*M2**4*
(t/z**2)*Q**6/(360.0*M_bar**3))
103             cs3 = Gf**2/(8*3.1415926)*(X3*N3**2*(2-2*t/z+(t/z)**2)-X3*N3**2*M3*
(t/z**2)-R2_eff*X3*N3**2*M3*(2-2*t/z+(t/z)**2)*Q**2/(3*M_bar)+R2_eff*X3*N3**2*M3**2*
(t/z**2)*Q**2/(3*M_bar)+R2_eff**2*X3*N3**2*M3**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff**2*X3*N3**2*M3**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff*X3*N3**2*M3**2*(2-2*t/z+
(t/z)**2)*Q**4/(60.0*M_bar**2)-R4_eff*X3*N3**2*M3**3*(t/z**2)*Q**4/(60.0*M_bar**2)-
R2_eff*R4_eff*X3*N3**2*M3**3*(2-2*t/z+
(t/z)**2)*Q**6/(360.0*M_bar**3)+R2_eff*R4_eff*X3*N3**2*M3**4*
(t/z**2)*Q**6/(360.0*M_bar**3))
104             cs4 = Gf**2/(8*3.1415926)*(X4*N4**2*(2-2*t/z+(t/z)**2)-X4*N4**2*M4*
(t/z**2)-R2_eff*X4*N4**2*M4*(2-2*t/z+(t/z)**2)*Q**2/(3*M_bar)+R2_eff*X4*N4**2*M4**2*
(t/z**2)*Q**2/(3*M_bar)+R2_eff**2*X4*N4**2*M4**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff**2*X4*N4**2*M4**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff*X4*N4**2*M4**2*(2-2*t/z+
(t/z)**2)*Q**4/(60.0*M_bar**2)-R4_eff*X4*N4**2*M4**3*(t/z**2)*Q**4/(60.0*M_bar**2)-
R2_eff*R4_eff*X4*N4**2*M4**3*(2-2*t/z+

```



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(t/z)**2)*Q**6/(360.0*M_bar**3)+R2_eff*R4_eff*X4*N4**2*M4**4*
(t/z**2)*Q**6/(360.0*M_bar**3))
104      cs5 = Gf**2/(8*3.1415926)*(X5*N5**2*(2-2*t/z+(t/z)**2)-X5*N5**2*M5*
(t/z**2)-R2_eff*X5*N5**2*M5*(2-2*t/z+(t/z)**2)*Q**2/(3*M_bar)+R2_eff*X5*N5**2*M5**2*
(t/z**2)*Q**2/(3*M_bar)+R2_eff**2*X5*N5**2*M5**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff**2*X5*N5**2*M5**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff*X5*N5**2*M5**2*(2-2*t/z+
(t/z)**2)*Q**4/(60.0*M_bar**2)-R4_eff*X5*N5**2*M5**3*(t/z**2)*Q**4/(60.0*M_bar**2)-
R2_eff*R4_eff*X5*N5**2*M5**3*(2-2*t/z+
(t/z)**2)*Q**6/(360.0*M_bar**3)+R2_eff*R4_eff*X5*N5**2*M5**4*
(t/z**2)*Q**6/(360.0*M_bar**3))
105      cross_section = cs1 + cs2 + cs3 + cs4 + cs5
106
107      ## flux
108      fu = 16/m**4*(3*m*z**2-4*z**3)
109      fe = 96/m**4*(m*z**2-2*z**3)
110
111      ## The full integral,multiply by (hbar*c)^2 to normalize.
112      C = 10**11
113      f2 = (hbar*c)**2*C*fe*cross_section*(3.156e7)*M_detector/1000
114      return f2
115
116      ## The proton part for flux fu.
117      def integrand3(z):
118          # Calculate dN/dt numerically, variable z is the energy E.
119          Q = (2*(z**2)*t*M_bar/(z**2-z*t))**0.5/(hbar*c)
120          # Five parts of cross section.
121          cs1 = Gf**2/(8*3.1415926)*(X1*Z**2*(2-2*t/z+(t/z)**2)-X1*Z**2*M1*
(t/z**2)-R2_eff_p*X1*Z**2*Z*(2-2*t/z+(t/z)**2)*Q**2/(3*M_bar)+R2_eff_p*X1*Z**2*M1**2*
(t/z**2)*Q**2/(3*M_bar)+R2_eff_p**2*X1*Z**2*M1**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff_p**2*X1*Z**2*M1**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff_p*X1*Z**2*M1**2*(2-2*t/z+
(t/z)**2)*Q**4/(60.0*M_bar**2)-R4_eff_p*X1*Z**2*M1**3*(t/z**2)*Q**4/(60.0*M_bar**2)-
R2_eff_p*R4_eff_p*X1*Z**2*M1**3*(2-2*t/z+
(t/z)**2)*Q**6/(360.0*M_bar**3)+R2_eff_p*R4_eff_p*X1*Z**2*M1**4*
(t/z**2)*Q**6/(360.0*M_bar**3))
122          cs2 = Gf**2/(8*3.1415926)*(X2*Z**2*(2-2*t/z+(t/z)**2)-X2*Z**2*M2*
(t/z**2)-R2_eff_p*X2*Z**2*Z*(2-2*t/z+(t/z)**2)*Q**2/(3*M_bar)+R2_eff_p*X2*Z**2*M2**2*
(t/z**2)*Q**2/(3*M_bar)+R2_eff_p**2*X2*Z**2*M2**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff_p**2*X2*Z**2*M2**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff_p*X2*Z**2*M2**2*(2-2*t/z+
(t/z)**2)*Q**4/(60.0*M_bar**2)-R4_eff_p*X2*Z**2*M2**3*(t/z**2)*Q**4/(60.0*M_bar**2)-
R2_eff_p*R4_eff_p*X2*Z**2*M2**3*(2-2*t/z+
(t/z)**2)*Q**6/(360.0*M_bar**3)+R2_eff_p*R4_eff_p*X2*Z**2*M2**4*
(t/z**2)*Q**6/(360.0*M_bar**3))
123          cs3 = Gf**2/(8*3.1415926)*(X3*Z**2*(2-2*t/z+(t/z)**2)-X3*Z**2*M3*
(t/z**2)-R2_eff_p*X3*Z**2*Z*(2-2*t/z+(t/z)**2)*Q**2/(3*M_bar)+R2_eff_p*X3*Z**2*M3**2*
(t/z**2)*Q**2/(3*M_bar)+R2_eff_p**2*X3*Z**2*M3**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff_p**2*X3*Z**2*M3**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff_p*X3*Z**2*M3**2*(2-2*t/z+
(t/z)**2)*Q**4/(60.0*M_bar**2)-R4_eff_p*X3*Z**2*M3**3*(t/z**2)*Q**4/(60.0*M_bar**2)-
R2_eff_p*R4_eff_p*X3*Z**2*M3**3*(2-2*t/z+
(t/z)**2)*Q**6/(360.0*M_bar**3)+R2_eff_p*R4_eff_p*X3*Z**2*M3**4*
(t/z**2)*Q**6/(360.0*M_bar**3))
124          cs4 = Gf**2/(8*3.1415926)*(X4*Z**2*(2-2*t/z+(t/z)**2)-X4*Z**2*M4*
(t/z**2)-R2_eff_p*X4*Z**2*Z*(2-2*t/z+(t/z)**2)*Q**2/(3*M_bar)+R2_eff_p*X4*Z**2*M4**2*
(t/z**2)*Q**2/(3*M_bar)+R2_eff_p**2*X4*Z**2*M4**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff_p**2*X4*Z**2*M4**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff_p*X4*Z**2*M4**2*(2-2*t/z+
(t/z)**2)*Q**4/(60.0*M_bar**2)-R4_eff_p*X4*Z**2*M4**3*(t/z**2)*Q**4/(60.0*M_bar**2)-

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```

R2_eff_p*R4_eff_p*X4*Z**2*M4**3*(2-2*t/z+
(t/z)**2)*Q**6/(360.0*M_bar**3)+R2_eff_p*R4_eff_p*X4*Z**2*M4**4*
(t/z**2)*Q**6/(360.0*M_bar**3))
125     cs5 = Gf**2/(8*3.1415926)*(X5*Z**2*(2-2*t/z+(t/z)**2)-X5*Z**2*M5*
(t/z**2)-R2_eff_p*X5*Z**2*Z*(2-2*t/z+(t/z)**2)*Q**2/(3*M_bar)+R2_eff_p*X5*Z**2*M5**2*
(t/z**2)*Q**2/(3*M_bar)+R2_eff_p**2*X5*Z**2*M5**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff_p**2*X5*Z**2*M5**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff_p*X5*Z**2*M5**2*(2-2*t/z+
(t/z)**2)*Q**4/(60.0*M_bar**2)-R4_eff_p*X5*Z**2*M5**3*(t/z**2)*Q**4/(60.0*M_bar**2)-
R2_eff_p*R4_eff_p*X5*Z**2*M5**3*(2-2*t/z+
(t/z)**2)*Q**6/(360.0*M_bar**3)+R2_eff_p*R4_eff_p*X5*Z**2*M5**4*
(t/z**2)*Q**6/(360.0*M_bar**3))
126     cross_section = alpha**2*(cs1 + cs2 + cs3 + cs4 + cs5)
127
128     ## flux
129     fu = 16/m**4*(3*m*z**2-4*z**3)
130     fe = 96/m**4*(m*z**2-2*z**3)
131
132     ## The full integral,multiply by (hbar*c)^2 to normalize.
133     C = 10**11
134     f3 = (hbar*c)**2*C*fu*cross_section*(3.156e7)*M_detector/1000
135     return f3
136
137     ## The proton part for flux fe.
138     def integrand4(z):
139         # Calculate dN/dt numerically, variable z is the energy E.
140         Q = (2*(z**2)*t*M_bar/(z**2-z*t))**0.5/(hbar*c)
141         # Five parts of cross section.
142         cs1 = Gf**2/(8*3.1415926)*(X1*Z**2*(2-2*t/z+(t/z)**2)-X1*Z**2*M1*
(t/z**2)-R2_eff_p*X1*Z**2*Z*(2-2*t/z+(t/z)**2)*Q**2/(3*M_bar)+R2_eff_p*X1*Z**2*M1**2*
(t/z**2)*Q**2/(3*M_bar)+R2_eff_p**2*X1*Z**2*M1**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff_p**2*X1*Z**2*M1**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff_p*X1*Z**2*M1**2*(2-2*t/z+
(t/z)**2)*Q**4/(60.0*M_bar**2)-R4_eff_p*X1*Z**2*M1**3*(t/z**2)*Q**4/(60.0*M_bar**2)-
R2_eff_p*R4_eff_p*X1*Z**2*M1**3*(2-2*t/z+
(t/z)**2)*Q**6/(360.0*M_bar**3)+R2_eff_p*R4_eff_p*X1*Z**2*M1**4*
(t/z**2)*Q**6/(360.0*M_bar**3))
143         cs2 = Gf**2/(8*3.1415926)*(X2*Z**2*(2-2*t/z+(t/z)**2)-X2*Z**2*M2*
(t/z**2)-R2_eff_p*X2*Z**2*Z*(2-2*t/z+(t/z)**2)*Q**2/(3*M_bar)+R2_eff_p*X2*Z**2*M2**2*
(t/z**2)*Q**2/(3*M_bar)+R2_eff_p**2*X2*Z**2*M2**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff_p**2*X2*Z**2*M2**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff_p*X2*Z**2*M2**2*(2-2*t/z+
(t/z)**2)*Q**4/(60.0*M_bar**2)-R4_eff_p*X2*Z**2*M2**3*(t/z**2)*Q**4/(60.0*M_bar**2)-
R2_eff_p*R4_eff_p*X2*Z**2*M2**3*(2-2*t/z+
(t/z)**2)*Q**6/(360.0*M_bar**3)+R2_eff_p*R4_eff_p*X2*Z**2*M2**4*
(t/z**2)*Q**6/(360.0*M_bar**3))
144         cs3 = Gf**2/(8*3.1415926)*(X3*Z**2*(2-2*t/z+(t/z)**2)-X3*Z**2*M3*
(t/z**2)-R2_eff_p*X3*Z**2*Z*(2-2*t/z+(t/z)**2)*Q**2/(3*M_bar)+R2_eff_p*X3*Z**2*M3**2*
(t/z**2)*Q**2/(3*M_bar)+R2_eff_p**2*X3*Z**2*M3**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff_p**2*X3*Z**2*M3**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff_p*X3*Z**2*M3**2*(2-2*t/z+
(t/z)**2)*Q**4/(60.0*M_bar**2)-R4_eff_p*X3*Z**2*M3**3*(t/z**2)*Q**4/(60.0*M_bar**2)-
R2_eff_p*R4_eff_p*X3*Z**2*M3**3*(2-2*t/z+
(t/z)**2)*Q**6/(360.0*M_bar**3)+R2_eff_p*R4_eff_p*X3*Z**2*M3**4*
(t/z**2)*Q**6/(360.0*M_bar**3))
145         cs4 = Gf**2/(8*3.1415926)*(X4*Z**2*(2-2*t/z+(t/z)**2)-X4*Z**2*M4*
(t/z**2)-R2_eff_p*X4*Z**2*Z*(2-2*t/z+(t/z)**2)*Q**2/(3*M_bar)+R2_eff_p*X4*Z**2*M4**2*
(t/z**2)*Q**2/(3*M_bar)+R2_eff_p**2*X4*Z**2*M4**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff_p**2*X4*Z**2*M4**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff_p*X4*Z**2*M4**2*(2-2*t/z+

```

```

(t/z)**2)*Q**4/(60.0*M_bar**2)-R4_eff_p*X4*Z**2*M4**3*(t/z**2)*Q**4/(60.0*M_bar**2)-
R2_eff_p*R4_eff_p*X4*Z**2*M4**3*(2-2*t/z+
(t/z)**2)*Q**6/(360.0*M_bar**3)+R2_eff_p*R4_eff_p*X4*Z**2*M4**4*
(t/z**2)*Q**6/(360.0*M_bar**3))
146     cs5 = Gf**2/(8*3.1415926)*(X5*Z**2*(2-2*t/z+(t/z)**2)-X5*Z**2*M5*
(t/z**2)-R2_eff_p*X5*Z**2*Z*(2-2*t/z+(t/z)**2)*Q**2/(3*M_bar)+R2_eff_p*X5*Z**2*M5**2*
(t/z**2)*Q**2/(3*M_bar)+R2_eff_p**2*X5*Z**2*M5**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff_p**2*X5*Z**2*M5**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff_p*X5*Z**2*M5**2*(2-2*t/z+
(t/z)**2)*Q**4/(60.0*M_bar**2)-R4_eff_p*X5*Z**2*M5**3*(t/z**2)*Q**4/(60.0*M_bar**2)-
R2_eff_p*R4_eff_p*X5*Z**2*M5**3*(2-2*t/z+
(t/z)**2)*Q**6/(360.0*M_bar**3)+R2_eff_p*R4_eff_p*X5*Z**2*M5**4*
(t/z**2)*Q**6/(360.0*M_bar**3))
147     cross_section = alpha**2*(cs1 + cs2 + cs3 + cs4 + cs5)
148
149     ## flux
150     fu = 16/m**4*(3*m*z**2-4*z**3)
151     fe = 96/m**4*(m*z**2-2*z**3)
152
153     ## The full integral,multiply by (hbar*c)^2 to normalize.
154     C = 10**11
155     f4 = (hbar*c)**2*C*fe*cross_section*(3.156e7)*M_detector/1000
156     return f4
157     ## The Cross term for flux fu and fe.
158     def integrand5(z,R2,R4):
159         R2_eff = R2
160         R4_eff = R4
161         # Calculate dN/dt numerically, variable z is the energy E.
162         Q = (2*(z**2)*t*M_bar/(z**2-z*t))**0.5/(hbar*c)
163         # Five parts of cross section.
164         cs1 = Gf**2/(8*3.1415926)*(X1*N1*Z*(2-2*t/z+(t/z)**2)-X1*N1*Z*M1*
(t/z**2)-R2_eff*X1*N1**2*M1*(2-2*t/z+(t/z)**2)*Q**2/(6*M_bar)+R2_eff*X1*N1**2*M1**2*
(t/z**2)*Q**2/(6*M_bar)-R2_eff_p*X1*Z**2*M1*(2-2*t/z+
(t/z)**2)*Q**2/(6*M_bar)+R2_eff_p*X1*Z**2*M1**2*
(t/z**2)*Q**2/(6*M_bar)+R2_eff*R2_eff_p*X1*N1*Z*M1**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff*R2_eff_p*X1*N1*Z*M1**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff*X1*N1**2*M1**2*(2-2*t/z+
(t/z)**2)*Q**4/(120.0*M_bar**2)-R4_eff*X1*N1**2*M1**3*
(t/z**2)*Q**4/(120.0*M_bar**2)+R4_eff_p*X1*Z**2*M1**2*(2-2*t/z+
(t/z)**2)*Q**4/(120.0*M_bar**2)-R4_eff_p*X1*Z**2*M1**3*
(t/z**2)*Q**4/(120.0*M_bar**2)-R2_eff_p*R4_eff*X1*N1*Z*M1**3*(2-2*t/z+
(t/z)**2)*Q**6/(720.0*M_bar**3)+R2_eff_p*R4_eff*X1*N1*Z*M1**4*
(t/z**2)*Q**6/(720.0*M_bar**3)-R2_eff*R4_eff_p*X1*N1*Z*M1**3*(2-2*t/z+
(t/z)**2)*Q**6/(720.0*M_bar**3)+R2_eff_p*R4_eff_p*X1*N1*Z*M1**4*
(t/z**2)*Q**6/(720.0*M_bar**3))
165         cs2 = Gf**2/(8*3.1415926)*(X2*N2*Z*(2-2*t/z+(t/z)**2)-X2*N2*Z*M2*
(t/z**2)-R2_eff*X2*N2**2*M2*(2-2*t/z+(t/z)**2)*Q**2/(6*M_bar)+R2_eff*X2*N2**2*M2**2*
(t/z**2)*Q**2/(6*M_bar)-R2_eff_p*X2*Z**2*M2*(2-2*t/z+
(t/z)**2)*Q**2/(6*M_bar)+R2_eff_p*X2*Z**2*M2**2*
(t/z**2)*Q**2/(6*M_bar)+R2_eff*R2_eff_p*X2*N2*Z*M2**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff*R2_eff_p*X2*N2*Z*M2**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff*X2*N2**2*M2**2*(2-2*t/z+
(t/z)**2)*Q**4/(120.0*M_bar**2)-R4_eff*X2*N2**2*M2**3*
(t/z**2)*Q**4/(120.0*M_bar**2)+R4_eff_p*X2*Z**2*M2**2*(2-2*t/z+
(t/z)**2)*Q**4/(120.0*M_bar**2)-R4_eff_p*X2*Z**2*M2**3*
(t/z**2)*Q**4/(120.0*M_bar**2)-R2_eff_p*R4_eff*X2*N2*Z*M2**3*(2-2*t/z+
(t/z)**2)*Q**6/(720.0*M_bar**3)+R2_eff_p*R4_eff*X2*N2*Z*M2**4*
(t/z**2)*Q**6/(720.0*M_bar**3)-R2_eff*R4_eff_p*X2*N2*Z*M2**3*(2-2*t/z+
(t/z)**2)*Q**6/(720.0*M_bar**3)+R2_eff_p*R4_eff_p*X2*N2*Z*M2**4*
(t/z**2)*Q**6/(720.0*M_bar**3))

```

```

166     cs3 = Gf**2/(8*3.1415926)*(X3*N3*Z*(2-2*t/z+(t/z)**2)-X3*N3*Z*M3*
(t/z**2)-R2_eff*X3*N3**2*M3*(2-2*t/z+(t/z)**2)*Q**2/(6*M_bar)+R2_eff*X3*N3**2*M3**2*
(t/z**2)*Q**2/(6*M_bar)-R2_eff_p*X3*Z**2*M3*(2-2*t/z+
(t/z)**2)*Q**2/(6*M_bar)+R2_eff_p*X3*Z**2*M3**2*
(t/z**2)*Q**2/(6*M_bar)+R2_eff*R2_eff_p*X3*N3*Z*M3**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff*R2_eff_p*X3*N3*Z*M3**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff*X3*N3**2*M3**2*(2-2*t/z+
(t/z)**2)*Q**4/(120.0*M_bar**2)-R4_eff*X3*N3**2*M3**3*
(t/z**2)*Q**4/(120.0*M_bar**2)+R4_eff_p*X3*Z**2*M3**2*(2-2*t/z+
(t/z)**2)*Q**4/(120.0*M_bar**2)-R4_eff_p*X3*Z**2*M3**3*
(t/z**2)*Q**4/(120.0*M_bar**2)-R2_eff_p*R4_eff*X3*N3*Z*M3**3*(2-2*t/z+
(t/z)**2)*Q**6/(720.0*M_bar**3)+R2_eff_p*R4_eff*X3*N3*Z*M3**4*
(t/z**2)*Q**6/(720.0*M_bar**3)-R2_eff*R4_eff_p*X3*N3*Z*M3**3*(2-2*t/z+
(t/z)**2)*Q**6/(720.0*M_bar**3)+R2_eff*R4_eff_p*X3*N3*Z*M3**4*
(t/z**2)*Q**6/(720.0*M_bar**3))
167     cs4 = Gf**2/(8*3.1415926)*(X4*N4*Z*(2-2*t/z+(t/z)**2)-X4*N4*Z*M4*
(t/z**2)-R2_eff*X4*N4**2*M4*(2-2*t/z+(t/z)**2)*Q**2/(6*M_bar)+R2_eff*X4*N4**2*M4**2*
(t/z**2)*Q**2/(6*M_bar)-R2_eff_p*X4*Z**2*M4*(2-2*t/z+
(t/z)**2)*Q**2/(6*M_bar)+R2_eff_p*X4*Z**2*M4**2*
(t/z**2)*Q**2/(6*M_bar)+R2_eff*R2_eff_p*X4*N4*Z*M4**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff*R2_eff_p*X4*N4*Z*M4**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff*X4*N4**2*M4**2*(2-2*t/z+
(t/z)**2)*Q**4/(120.0*M_bar**2)-R4_eff*X4*N4**2*M4**3*
(t/z**2)*Q**4/(120.0*M_bar**2)+R4_eff_p*X4*Z**2*M4**2*(2-2*t/z+
(t/z)**2)*Q**4/(120.0*M_bar**2)-R4_eff_p*X4*Z**2*M4**3*
(t/z**2)*Q**4/(120.0*M_bar**2)-R2_eff_p*R4_eff*X4*N4*Z*M4**3*(2-2*t/z+
(t/z)**2)*Q**6/(720.0*M_bar**3)+R2_eff_p*R4_eff*X4*N4*Z*M4**4*
(t/z**2)*Q**6/(720.0*M_bar**3)-R2_eff*R4_eff_p*X4*N4*Z*M4**3*(2-2*t/z+
(t/z)**2)*Q**6/(720.0*M_bar**3)+R2_eff*R4_eff_p*X4*N4*Z*M4**4*
(t/z**2)*Q**6/(720.0*M_bar**3))
168     cs5 = Gf**2/(8*3.1415926)*(X5*N5*Z*(2-2*t/z+(t/z)**2)-X5*N5*Z*M5*
(t/z**2)-R2_eff*X5*N5**2*M5*(2-2*t/z+(t/z)**2)*Q**2/(6*M_bar)+R2_eff*X5*N5**2*M5**2*
(t/z**2)*Q**2/(6*M_bar)-R2_eff_p*X5*Z**2*M5*(2-2*t/z+
(t/z)**2)*Q**2/(6*M_bar)+R2_eff_p*X5*Z**2*M5**2*
(t/z**2)*Q**2/(6*M_bar)+R2_eff*R2_eff_p*X5*N5*Z*M5**2*(2-2*t/z+
(t/z)**2)*Q**4/(36.0*M_bar**2)-R2_eff*R2_eff_p*X5*N5*Z*M5**3*
(t/z**2)*Q**4/(36.0*M_bar**2)+R4_eff*X5*N5**2*M5**2*(2-2*t/z+
(t/z)**2)*Q**4/(120.0*M_bar**2)-R4_eff*X5*N5**2*M5**3*
(t/z**2)*Q**4/(120.0*M_bar**2)+R4_eff_p*X5*Z**2*M5**2*(2-2*t/z+
(t/z)**2)*Q**4/(120.0*M_bar**2)-R4_eff_p*X5*Z**2*M5**3*
(t/z**2)*Q**4/(120.0*M_bar**2)-R2_eff_p*R4_eff*X5*N5*Z*M5**3*(2-2*t/z+
(t/z)**2)*Q**6/(720.0*M_bar**3)+R2_eff_p*R4_eff*X5*N5*Z*M5**4*
(t/z**2)*Q**6/(720.0*M_bar**3)-R2_eff*R4_eff_p*X5*N5*Z*M5**3*(2-2*t/z+
(t/z)**2)*Q**6/(720.0*M_bar**3)+R2_eff*R4_eff_p*X5*N5*Z*M5**4*
(t/z**2)*Q**6/(720.0*M_bar**3))
169     cross_section = -1*alpha**2*(cs1 + cs2 + cs3 + cs4 + cs5)
170
171     ## flux
172     fu = 16/m**4*(3*m*z**2-4*z**3)
173     fe = 96/m**4*(m*z**2-2*z**3)
174
175     ## The full integral,multiply by (hbar*c)^2 to normalize.
176     C = 10**11
177     f5 = (hbar*c)**2*C*(fu+fe)*cross_section*(3.156e7)*M_detector/1000
178     return f5
179
180     I1 = scipy.integrate.quad(integrand1,(t+sqrt(t**2+2*t*M_bar))/2,m/2,args=
(R2,R4))
181     I2 = scipy.integrate.quad(integrand2,(t+sqrt(t**2+2*t*M_bar))/2,m/2,args=
(R2,R4))

```

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182         I3 = scipy.integrate.quad(integrand3,(t+sqrt(t**2+2*t*M_bar))/2,m/2)
183         I4 = scipy.integrate.quad(integrand4,(t+sqrt(t**2+2*t*M_bar))/2,m/2)
184         I5 = scipy.integrate.quad(integrand5,(t+sqrt(t**2+2*t*M_bar))/2,m/2,args=
(R2,R4))
185         ssum = I1[0]+I2[0]+I3[0]+I4[0]+I5[0]
186         #print(ssum)
187         sum_bin = sum_bin + ssum
188         event.append(int(ssum))
189
190         event_bin.append(int(sum_bin))
191         #print(event)
192         return (event,event_bin)
193
194
195 # PART3: Plots and Exporting data.
196 #####
197 energy = np.linspace(1,80,num=80)
198 energy2 = np.linspace(5,80,num=16)
199 (event1,event_bin1) = dNdt(R2,R4)
200 # (event2,event_bin2) = dNdt(R2*0.9**2,R4*0.9**4)
201 # (event3,event_bin3) = dNdt(R2*0.8**2,R4*0.8**4)
202 # (event4,event_bin4) = dNdt(R2*1.1**2,R4*1.1**4)
203 # (event5,event_bin5) = dNdt(R2*1.2**2,R4*1.2**4)
204
205 # print("event=",event1)
206 # print("binned event=",event_bin1)
207
208 # table = pd.DataFrame({ 'Bin Range(keV)': energy, 'Events' : event1, 'Events (-10%)'
: event2, 'Events (-20%)' : event3, 'Events (+10%)' : event4, 'Events (+20%)' :
event5})
209 # table.to_csv('table_Ge_with_proton.csv', index=False, encoding='utf-8')
210
211
212 # #plt.plot(energy,event1)
213 # plt.bar(energy2,event_bin1,align='center',width=4)
214 # plt.xticks(energy2)
215 # plt.ylabel("Number of Counts")
216 # plt.xlabel("Energy/keV")
217 # plt.title("# of Events per (keV year tonne) for Ge")
218 # for i, v in enumerate(event_bin1):
219 #     plt.text(energy2[i] -0.25, v , str(v))
220 # plt.show()
221
222 # Bar Plot:
223 plt.rcParams['font.size'] = 10
224 plt.rcParams['axes.linewidth'] = 2
225 fig = plt.figure()
226 ax = fig.add_axes([0, 0, 1, 1])
227 # Edit the major and minor ticks of the x and y axes
228 ax.xaxis.set_tick_params(which='major', size=2, width=2, direction='in', top='on')
229 ax.yaxis.set_tick_params(which='major', size=5, width=2, direction='in', right='on')
230 ax.yaxis.set_tick_params(which='minor', size=2, width=2, direction='in', right='on')
231
232 # Set the axis limits
233 ax.set_xlim(0, 80)
234 ax.set_ylim(0, 10000)
235 # Edit the major and minor tick locations
236 ax.yaxis.set_major_locator(MultipleLocator(1000))
237 ax.yaxis.set_major_formatter(FormatStrFormatter('%d'))
238 ax.yaxis.set_minor_locator(AutoMinorLocator(2))

```



```

239
240 plt.bar(energy2,event_bin1,align='center',width=4)
241 plt.xticks(energy2)
242 plt.xticks(rotation=90)
243 plt.ylabel("Number of Counts (keV year tonne)")
244 plt.xlabel("Nucleus Recoil Energy (keV)")
245 for i, v in enumerate(event_bin1):
246     plt.text(energy2[i] -0.25, v , str(v))
247 plt.savefig('Scattering_Events_Ge_with_proton_binned.png', dpi=300,
transparent=False, bbox_inches='tight')
248 plt.show()
249
250
251 # ## Second Plot: Compare the difference between this and the one without proton.
252 # (energy0,event0) = np.loadtxt('table_Ge_data.csv', unpack=True, delimiter=',',
skiprows=1)
253 # diff = event1-event0
254 # plt.rcParams['font.size'] = 10
255 # plt.rcParams['axes.linewidth'] = 2
256 # fig = plt.figure()
257 # ax2 = fig.add_axes([0, 0, 1, 0.4])
258 # ax1 = fig.add_axes([0, 0.6, 1, 0.4])
259
260 # # # Set the axis limits
261 # ax1.set_xlim(0, 80)
262 # ax1.set_ylim(0, 2500)
263 # ax2.set_xlim(0, 80)
264 # ax2.set_ylim(-300, 80)
265 # # Edit the major and minor ticks of the x and y axes
266 # ax1.xaxis.set_tick_params(which='major', size=10, width=2, direction='in',
top='on')
267 # ax1.xaxis.set_tick_params(which='minor', size=7, width=2, direction='in', top='on')
268 # ax1.yaxis.set_tick_params(which='major', size=10, width=2, direction='in',
right='on')
269 # ax1.yaxis.set_tick_params(which='minor', size=7, width=2, direction='in',
right='on')
270 # ax2.xaxis.set_tick_params(which='major', size=10, width=2, direction='in',
top='on')
271 # ax2.xaxis.set_tick_params(which='minor', size=7, width=2, direction='in', top='on')
272 # ax2.yaxis.set_tick_params(which='major', size=10, width=2, direction='in',
right='on')
273 # ax2.yaxis.set_tick_params(which='minor', size=7, width=2, direction='in',
right='on')
274
275 # ax1.plot(energy,event1,label='With Proton and Cross Terms')
276 # ax1.plot(energy,event0,label='Without Proton and Cross Terms')
277 # ax2.plot(energy,diff)
278
279 # # Edit the major and minor tick locations
280 # ax1.xaxis.set_major_locator(MultipleLocator(20))
281 # ax1.xaxis.set_major_formatter(FormatStrFormatter('%d'))
282 # ax1.xaxis.set_minor_locator(AutoMinorLocator(5))
283 # ax1.yaxis.set_major_locator(MultipleLocator(500))
284 # ax1.yaxis.set_major_formatter(FormatStrFormatter('%d'))
285 # ax1.yaxis.set_minor_locator(AutoMinorLocator(5))
286 # ax2.xaxis.set_major_locator(MultipleLocator(20))
287 # ax2.xaxis.set_major_formatter(FormatStrFormatter('%d'))
288 # ax2.xaxis.set_minor_locator(AutoMinorLocator(5))
289 # ax2.yaxis.set_major_locator(MultipleLocator(50))
290 # ax2.yaxis.set_major_formatter(FormatStrFormatter('%d'))

```

```
291 # ax2.yaxis.set_minor_locator(AutoMinorLocator(5))
292 # # Add legend to plot
293 # ax1.legend(bbox_to_anchor=(0.95, 0.95),loc=1, frameon=False, fontsize=10)
294 # ax2.set_xlabel('Nucleus Recoil Energy (keV)', labelpad=10)
295 # ax2.set_ylabel('Difference in Events/(keV year tonne)', labelpad=10)
296 # ax1.set_ylabel('Events/(keV year tonne)', labelpad=10)
297 # plt.savefig('Scattering_Events_Ge_diff.png', dpi=300, transparent=False,
    bbox_inches='tight')
298 # plt.show()
299
300
```

```

1  ## Hongyong Zhang
2  ## Run Monte Carlo Simulations.
3  ## 05/19/2020
4  import matplotlib.pyplot as plt
5  import numpy as np
6  import matplotlib.pyplot as plt
7  import scipy.integrate
8  import pandas as pd
9  from numpy import cos, sin, exp, sqrt
10 from dNdt_Ge_binned import dNdt
11 import random
12 from scipy.stats import chisquare
13
14 ## Part1: Generate the database.
15 #####
16 R2 = (4.0495e-15)**2
17 R4 = (4.3765e-15)**4
18 events = []
19 ## Generate the standard events for reference first
20 standard_event = [R2,R4]
21 (result,result_binned) = dNdt(R2,R4)
22 standard_event = standard_event + result_binned
23 #print(standard_event)
24
25 ## Create database of events using R2 +/- 15% and R4 +/- 30% with 100 values for
    each.
26 for i in range(100):
27     R2 = (sqrt(R2)*0.3/100*(i+1)+sqrt(R2)*(1-0.15))**2
28     #print(i)
29     for p in range(100):
30         #print(p)
31         R4 = (R4**0.25*0.6/100*(i+1)+R4**0.25*(1-0.3))**4
32         event = [R2,R4]
33         (result,result_binned) = dNdt(R2,R4)
34         event = event + result_binned
35         events.append(event)
36         #print(event)
37 np.save("database.npy",events)
38 # To load the data: Use database = np.load("database.npy")
39 print(events)
40
41
42
43 ## Part2: Run the Monte Carlo simulation.
44 #####
45 R2 = (4.0495e-15)**2
46 R4 = (4.3765e-15)**4
47 minimal_R2=[]
48 minimal_R4=[]
49 for i in range(5000):
50
51     event_with_noise=[R2,R4]
52     ## generate noise for each term
53     for p in range(16):
54         noise = random.uniform(-1,1)*sqrt(standard_event[p+2]) + standard_event[p+2]
55         event_with_noise.append(noise)
56     chi_square=[]
57     ## chi square test
58     for q in range(len(events)):
59         expected = events[q][2:17]

```



```

60     obs = event_with_noise[2:17]
61     (chisq,p) = chisquare(obs,f_exp=expected)
62     chi_square.append(chisq)
63     ## find the minimum value and the corresponding R2 and R4
64     minimal_value_R2R4 = events[np.argmin(chi_square)][0:2]
65     minimal_R2.append(minimal_value_R2R4[0])
66     minimal_R4.append(minimal_value_R2R4[1])
67
68     print(minimal_value_R2R4)
69     ##print(event_with_noise)
70
71
72 ## Part3: Plots and Tables
73 #####
74 ## Scatter Plot
75 min_R2 = min(minimal_R2)
76 max_R2 = max(minimal_R2)
77 maxlim_R2 = max_R2 + abs((max_R2 - min_R2)*0.2)
78 minlim_R2 = min_R2 - abs((max_R2 - min_R2)*0.2)
79 min_R4 = min(minimal_R4)
80 max_R4 = max(minimal_R4)
81 maxlim_R4 = max_R4 + abs((max_R4 - min_R4)*0.2)
82 minlim_R4 = min_R4 - abs((max_R4 - min_R4)*0.2)
83 plt.scatter(minimal_R2,minimal_R4)
84 plt.xlim(minlim_R2,maxlim_R2)
85 plt.ylim(minlim_R4,maxlim_R4)
86 plt.show()
87
88 table = pd.DataFrame({ 'R2': minimal_R2, 'R4' : minimal_R4})
89 #table.to_csv('Monte_Carlo_new.csv', index=False, encoding='utf-8')

```

```

1  ## Hongyong Zhang
2  ## Plot the result of the Monte Carlo simulation.
3  ## 05/19/2020
4  import matplotlib as mpl
5  import matplotlib.pyplot as plt
6  from matplotlib import rc
7  from matplotlib.ticker import (MultipleLocator, FormatStrFormatter, AutoMinorLocator)
8  import numpy as np
9  import scipy.integrate
10 import pandas as pd
11 from numpy import cos, sin, exp, sqrt
12
13 (R2,R4) = np.loadtxt('Monte_Carlo_Ge.csv', unpack=True, delimiter=',', skiprows=1)
14
15 min_R2 = min(R2)
16 max_R2 = max(R2)
17 maxlim_R2 = max_R2 + abs((max_R2 - min_R2)*0.2)
18 minlim_R2 = min_R2 - abs((max_R2 - min_R2)*0.2)
19 min_R4 = min(R4)
20 max_R4 = max(R4)
21 maxlim_R4 = max_R4 + abs((max_R4 - min_R4)*0.2)
22 minlim_R4 = min_R4 - abs((max_R4 - min_R4)*0.2)
23 # Plot:
24 plt.rcParams['font.size'] = 10
25 plt.rcParams['axes.linewidth'] = 2
26 fig = plt.figure()
27 ax = fig.add_axes([0, 0, 1, 1])
28 # Edit the major and minor ticks of the x and y axes
29 ax.xaxis.set_tick_params(which='major', size=10, width=2, direction='in', top='on')
30 ax.xaxis.set_tick_params(which='minor', size=7, width=2, direction='in', top='on')
31 ax.yaxis.set_tick_params(which='major', size=10, width=2, direction='in', right='on')
32 ax.yaxis.set_tick_params(which='minor', size=7, width=2, direction='in', right='on')
33
34 ax.scatter(R2,R4)
35
36 # Set the axis limits
37 ax.set_xlim(minlim_R2, maxlim_R2)
38 ax.set_ylim(minlim_R4, maxlim_R4)
39 # Edit the major and minor tick locations
40 ax.xaxis.set_major_locator(MultipleLocator(0.1))
41 ax.xaxis.set_major_formatter(FormatStrFormatter('%.2f'))
42 ax.xaxis.set_minor_locator(AutoMinorLocator(5))
43 ax.yaxis.set_major_locator(MultipleLocator(0.4))
44 ax.yaxis.set_major_formatter(FormatStrFormatter('%.2f'))
45 ax.yaxis.set_minor_locator(AutoMinorLocator(5))
46
47 # Add legend to plot
48 ax.set_xlabel(r'$\langle R^2_n \rangle^{0.5}_{eff}$ (fm)$', labelpad=10)
49 ax.set_ylabel(r'$\langle R^4_n \rangle^{0.25}_{eff}$ (fm)$', labelpad=10)
50 plt.savefig('Monte_Carlo_Ge.png', dpi=300, transparent=False, bbox_inches='tight')
51 plt.show()
52

```